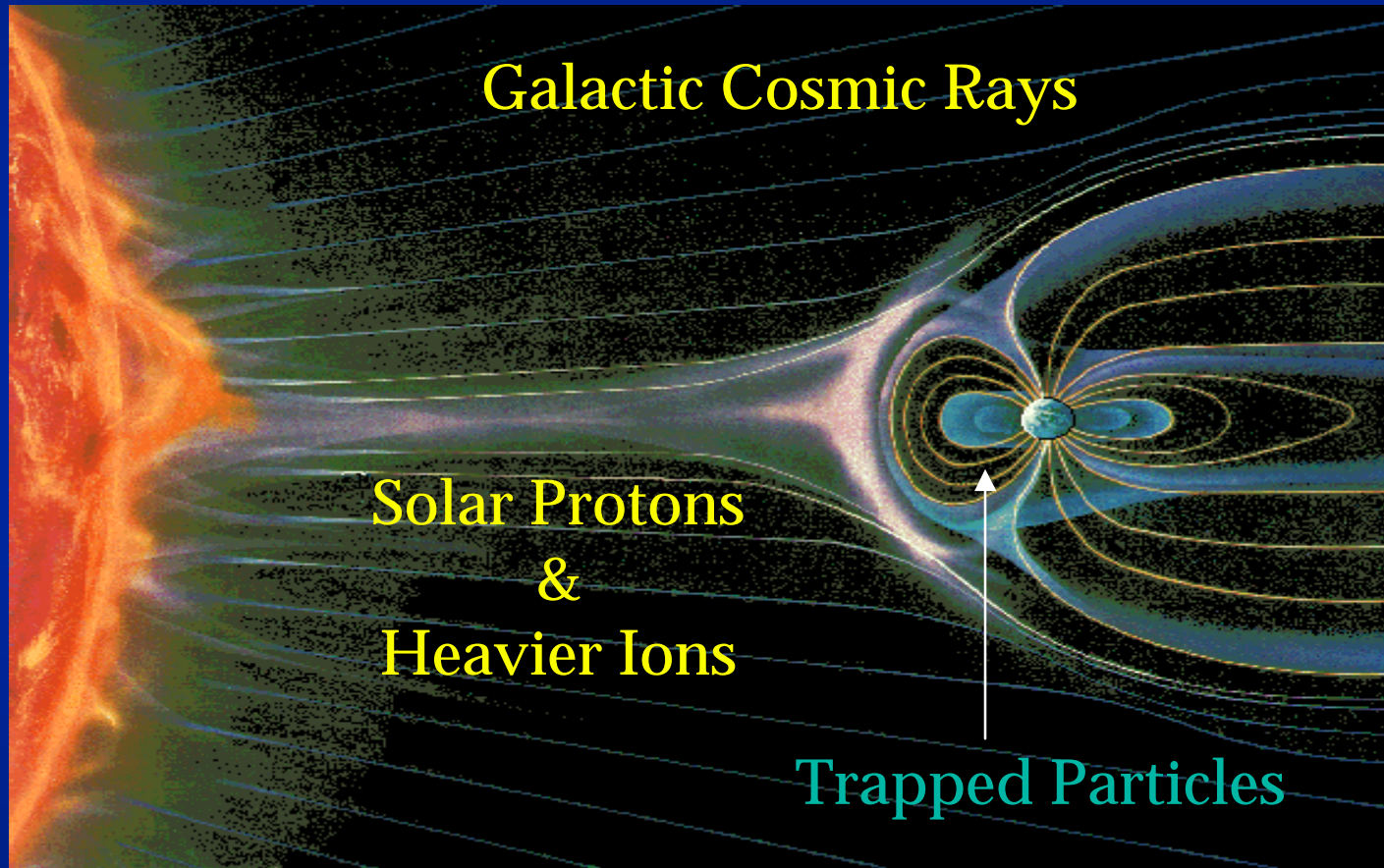
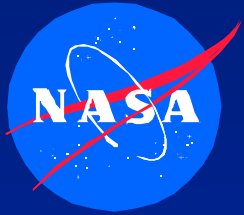


The Radiation Environment

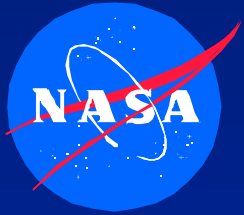


Nikkei Science, Inc. of Japan, by K. Endo



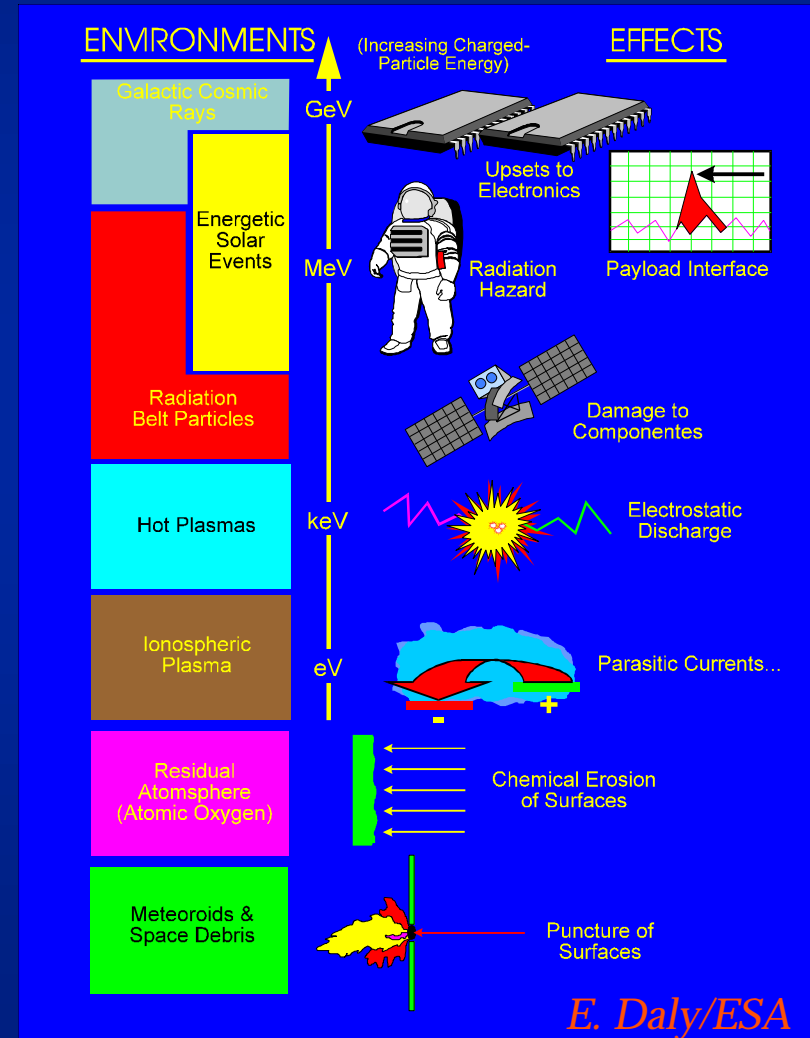
Components of the Natural Environment

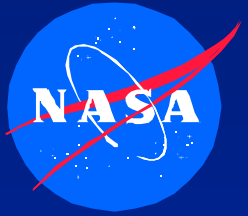
- ◆ Transient
 - » Galactic Cosmic Rays
 - Protons & Heavier Ions
 - » Solar Particle Events
 - Protons & Heavier Ions
- ◆ Trapped
 - » Electrons, Protons, & Heavier Ions
- ◆ Atmospheric & Terrestrial Secondaries
 - » Neutrons



Radiation & Effects

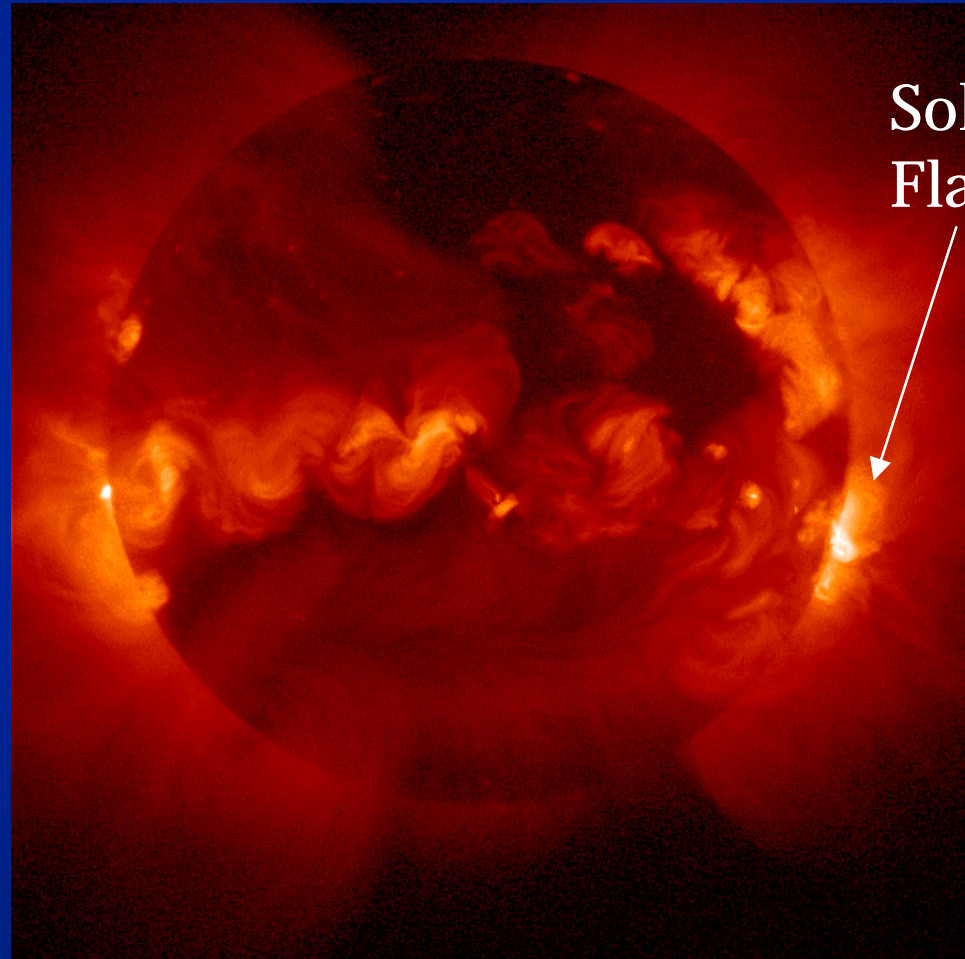
- ◆ Total Ionizing Dose
 - » Cumulative long term ionizing damage due to protons & electrons
- ◆ Displacement Damage
 - » Cumulative long term non-ionizing damage due to protons, electrons, & neutrons
- ◆ Single Event Effects
 - » Event caused by a single charged particle - heavy ions & protons for some devices



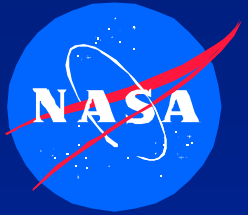


Sun

- ◆ Dominates the Radiation Environment
 - » Source
 - » Modulator
- ◆ Structure
 - » Photosphere
 - » Chromosphere
 - » Corona

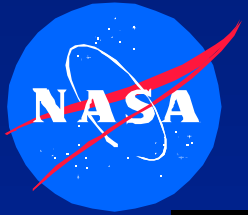


Solar
Flare



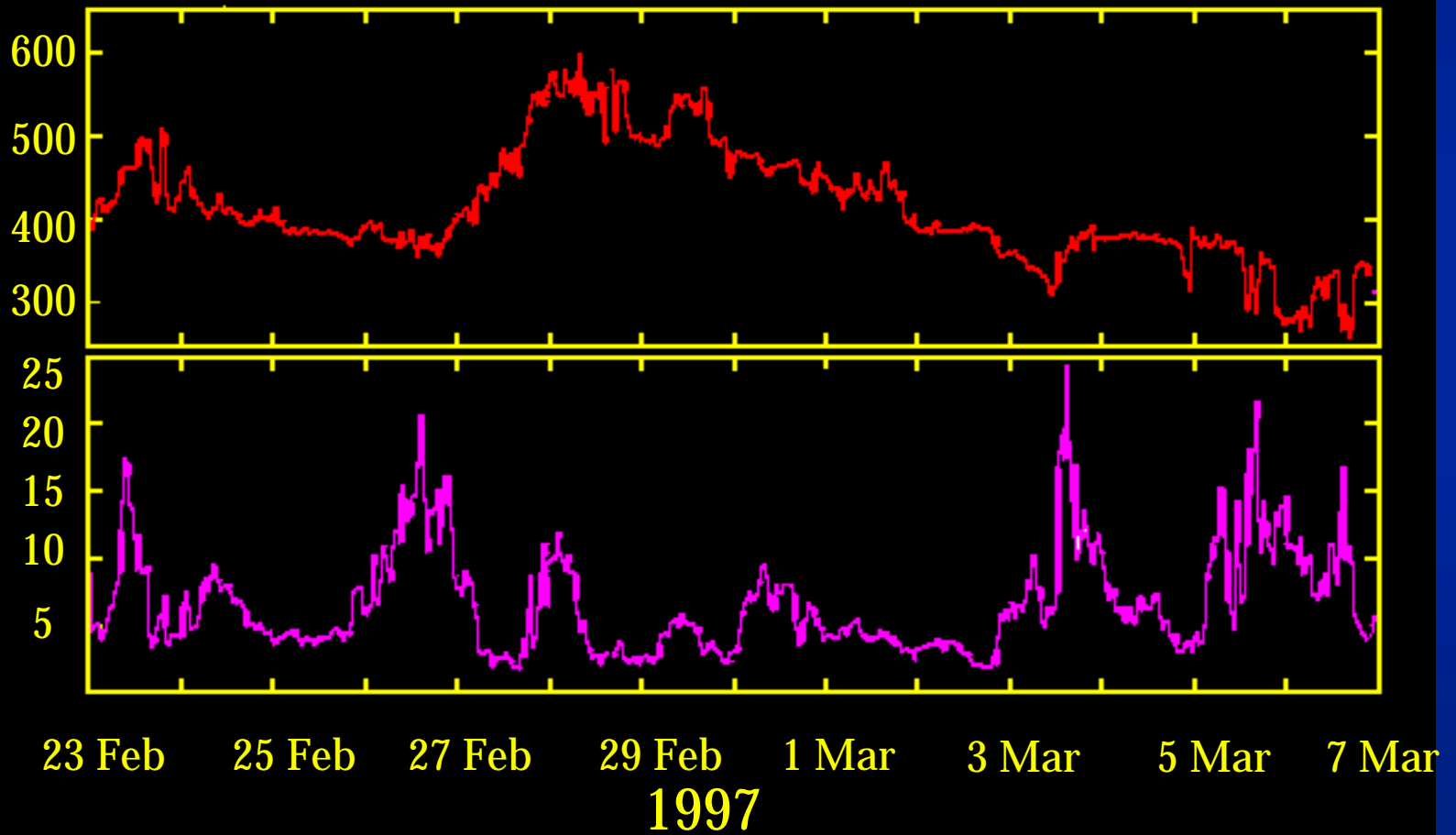
Solar Wind

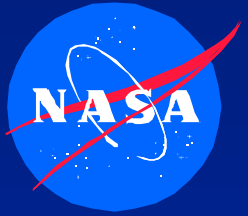
- ◆ Stream of Charged Particles from Sun's Corona
 - » Electrons
 - » Protons
 - » Heavy Ions
- } Density $\sim 1 - 30 / \text{cm}^3$
- ◆ Magnetized Plasma
 - ◆ Detected Out to 10 billion km from Earth by Pioneer 10
 - ◆ Velocity $\sim 300 - 900 \text{ km/s}$
 - ◆ Energy $\sim .5 - 2.0 \text{ keV/nuc}$



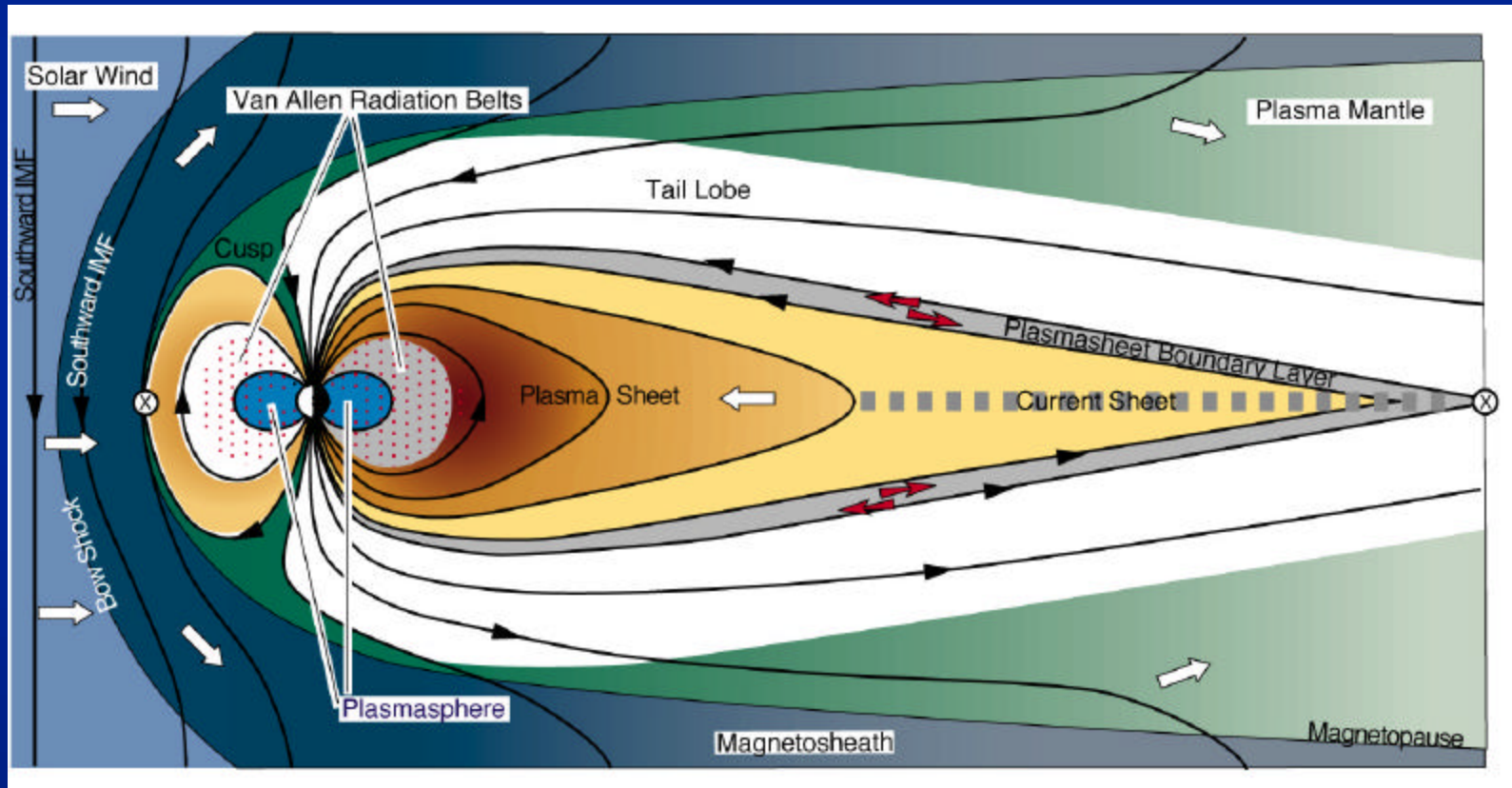
Solar Wind Density & Velocity

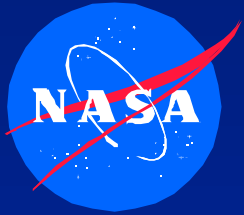
Density (#/cm³) Velocity (km/s)





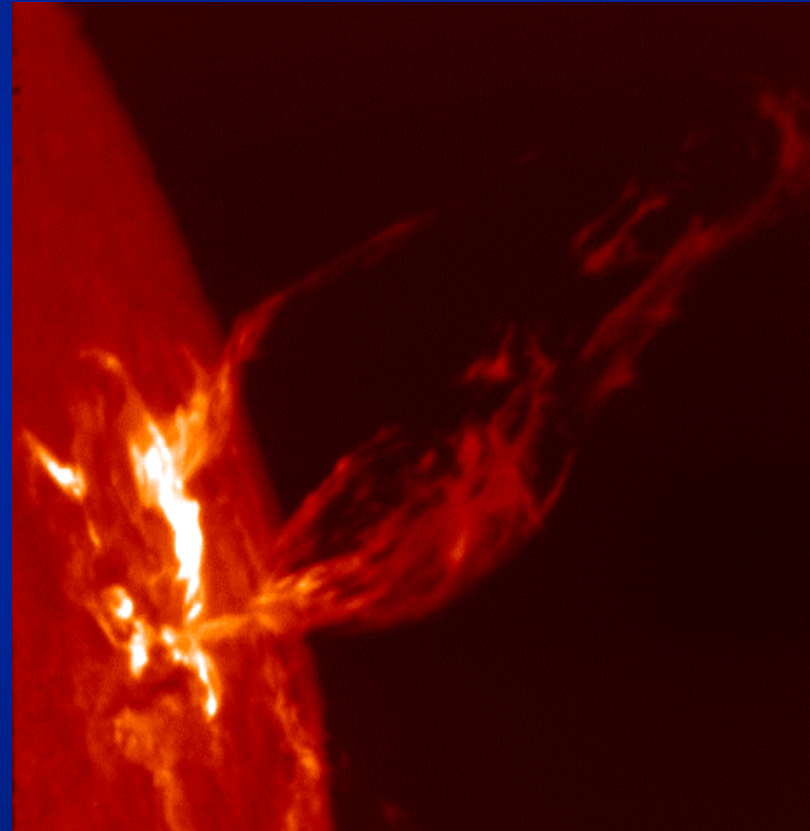
Magnetosphere



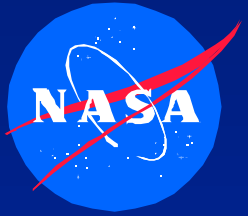


Coronal Mass Ejections

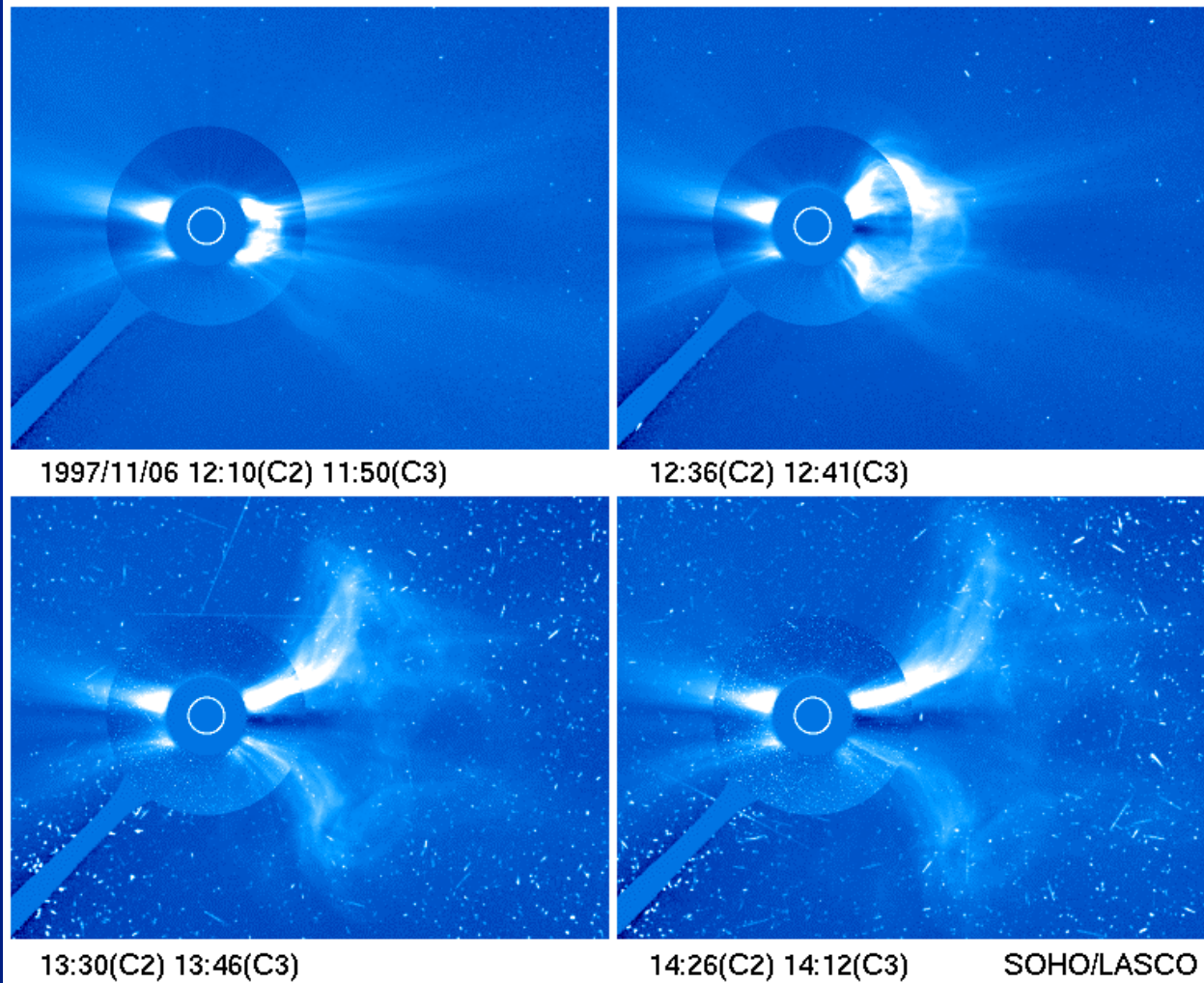
- ◆ Bubble of Gas & Magnetic Field
- ◆ Ejects $\sim 10^{17}$ grams of Matter
- ◆ Shock Wave Accelerates Particles to Millions of km/hr
- ◆ Associated with Magnetic Storms
- ◆ Proton Rich Solar Events

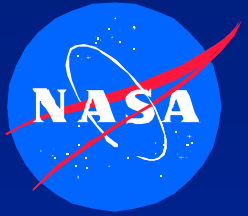


Holloman AFB/SOON



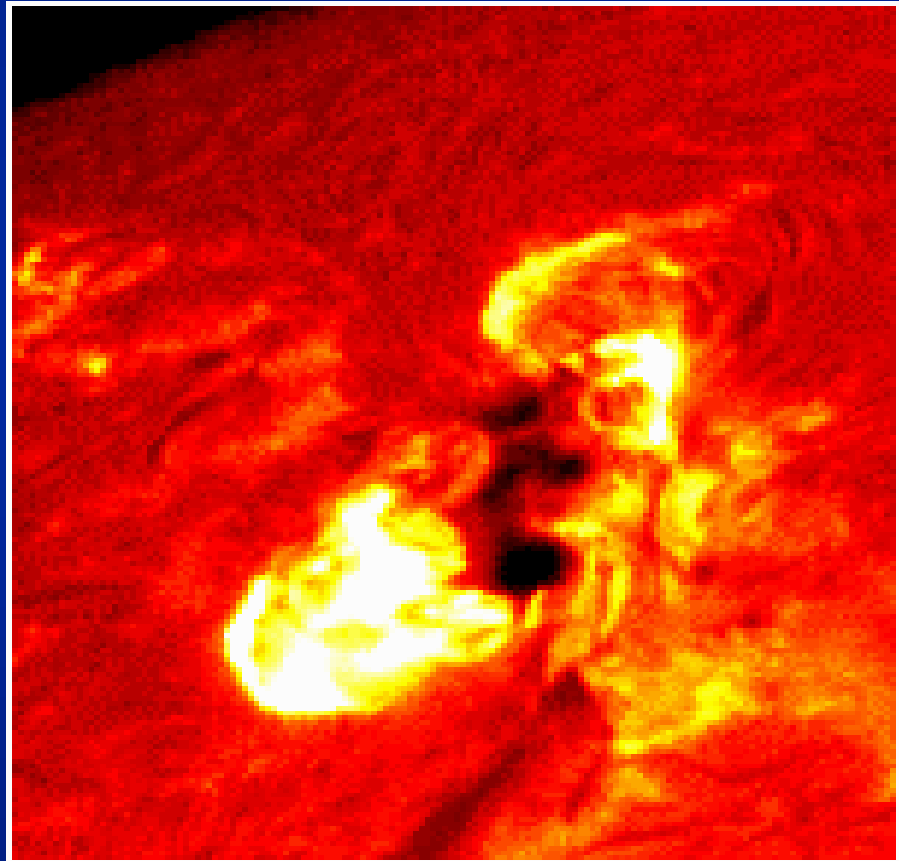
SOHO/LASCO During Solar Particle Event

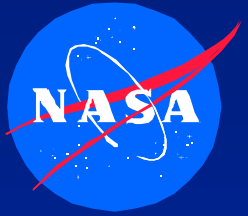




Solar Flares

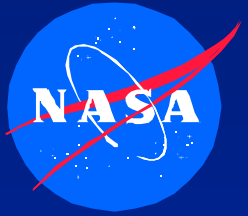
- ◆ Sudden Brightening Near Sunspots
- ◆ Solar System's Largest Explosive Events
- ◆ Particles Accelerated Directly by Event
- ◆ Heavy Ion Rich Solar Events





Magnetic Storms

- ◆ “Gusty” Solar Wind Disturbs the Current Systems in the Magnetosphere
- ◆ Major Storms Probably the Result of CMEs
 - » Must Be Pointed Toward Earth
 - » Strongest Arrive with Interplanetary Magnetic Field Oriented South
- ◆ Cause Increase in Rate & Intensity of Magnetic Sub-storms in the Geotail



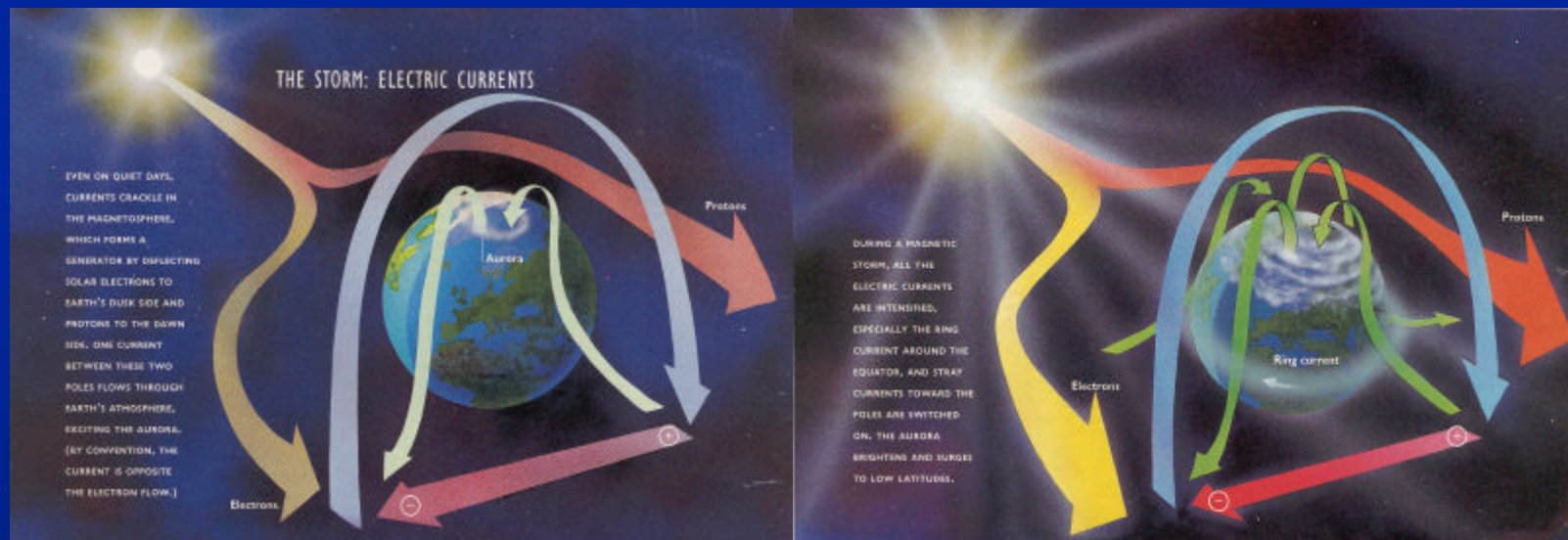
Electrical Currents

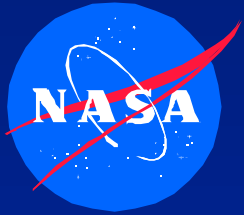
- ◆ Normal Conditions

- » Currents Present on Quiet Days

- ◆ Stormy Conditions

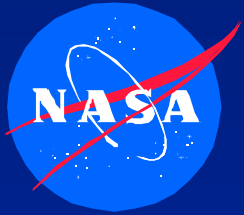
- » Intensified Currents





Effects of Solar Storms

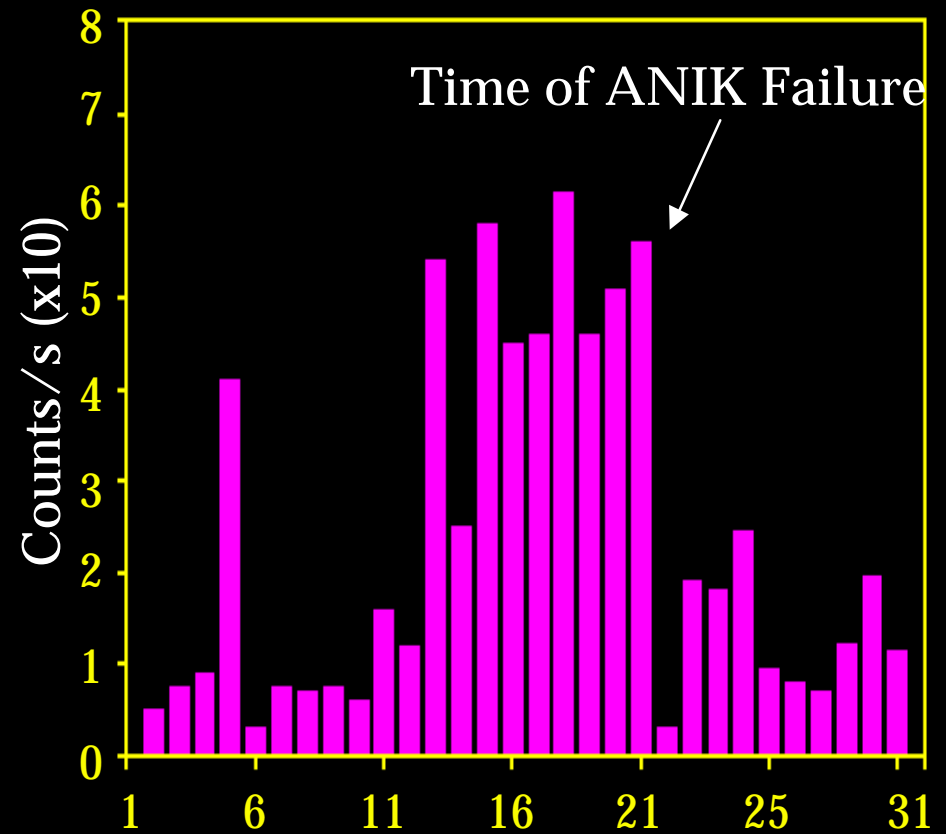
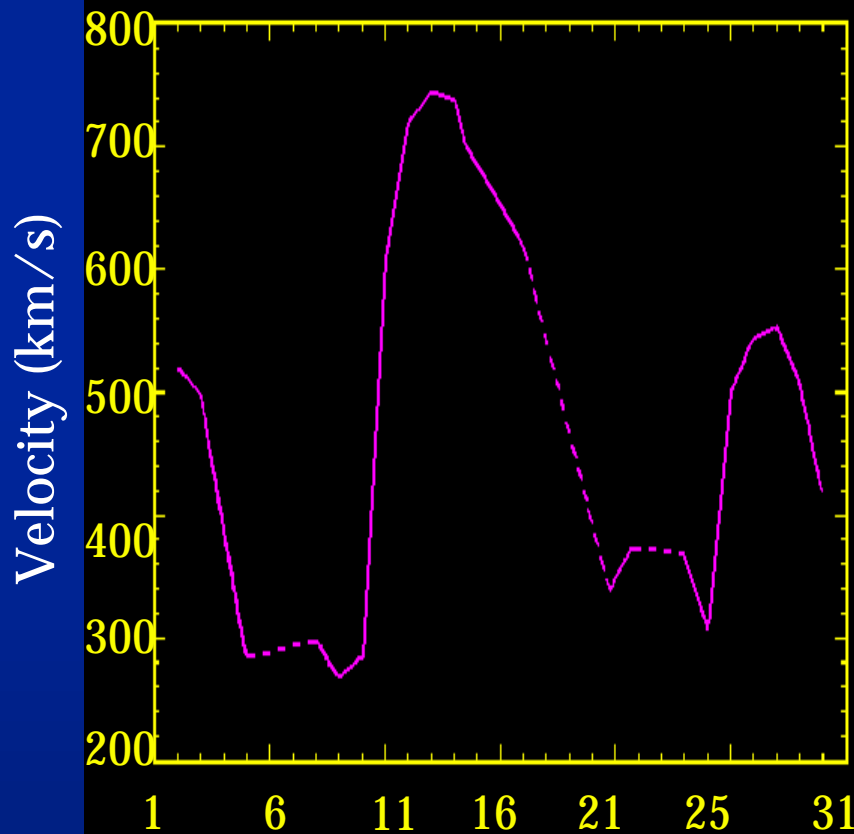
- ◆ Power Black-outs on Earth
- ◆ Block Some Radio Communication
 - » Disturbs electrically charged gases in the ionosphere
- ◆ Interfere with Cellular Phone Systems
 - » Ionospheric disturbances & satellite system failures
- ◆ Interfere with GPS Navigation (Ships & Airplanes)
- ◆ Trigger Phantom Commands on Spacecraft
- ◆ Increased Atmospheric Drag on Low Earth Orbit (LEO) Satellites
- ◆ Increased Protons & Heavy Ion Particle Counts
- ◆ “Pump Up” the Van Allen Belts

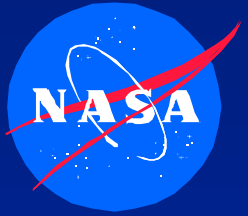


ANIK E1: Magnetic Storm

Solar Wind Velocity (IMP-8 MIT)

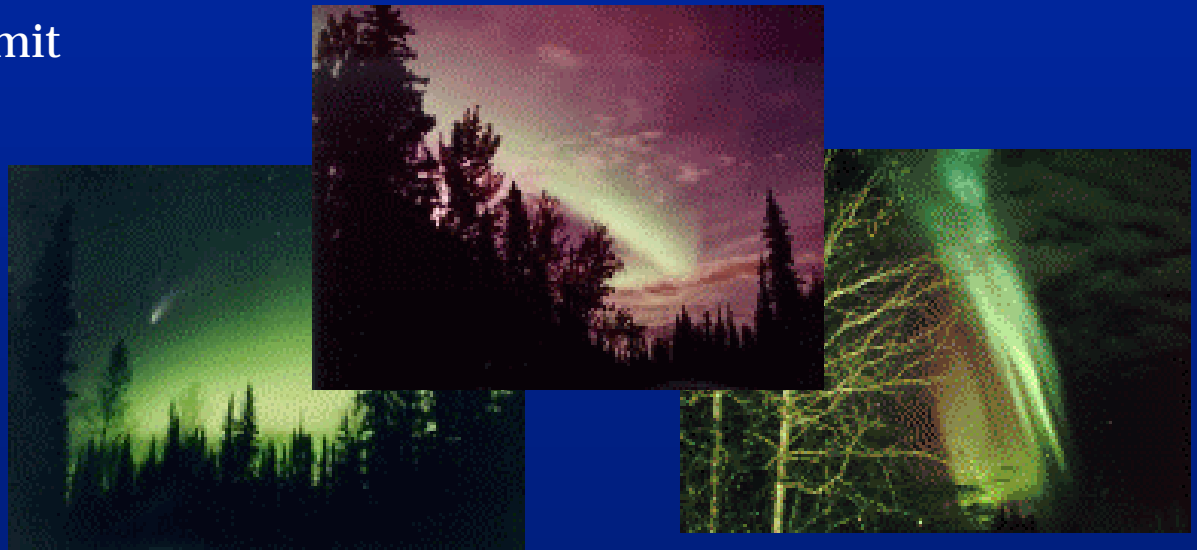
SAMPEX Electrons $E > 1$ MeV

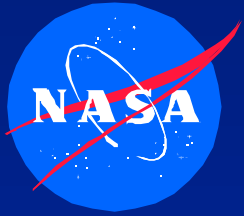




Aurora

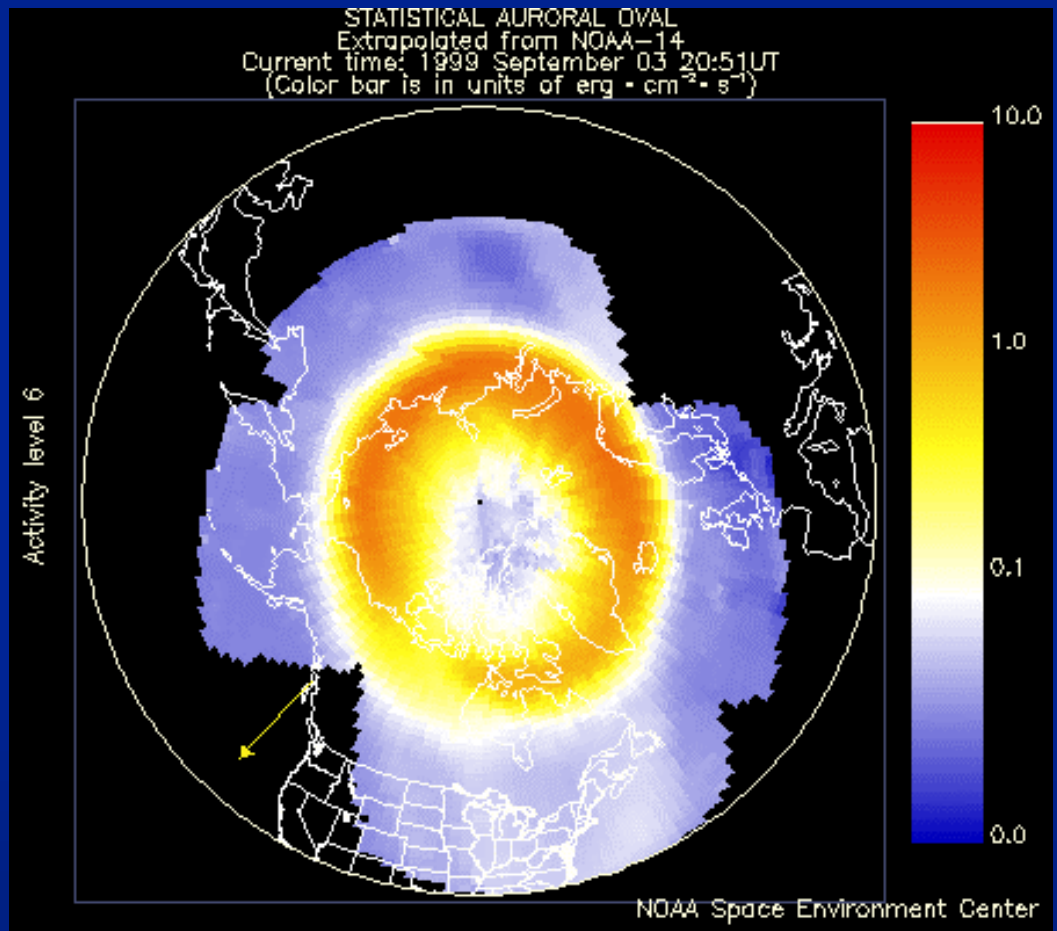
- ◆ Particles stream down on magnetic field lines from the geomagnetic tail forming an auroral belt
- ◆ Electrons collide with atmospheric gases
- ◆ Electrons give energy to atoms and molecules which emit energy as light
- ◆ Oxygen ---> Green
- ◆ Nitrogen ---> Red

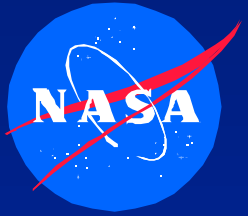




Aurora Borealis

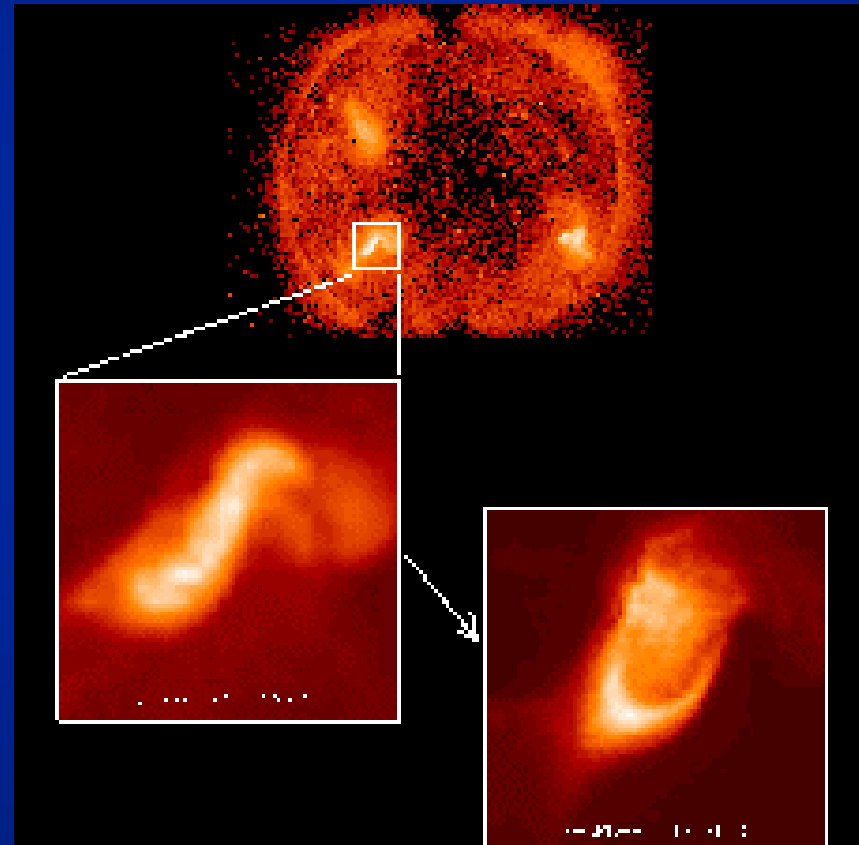
- ◆ Northern lights oval as measured by NOAA-14
- ◆ Centered on Magnetic Pole

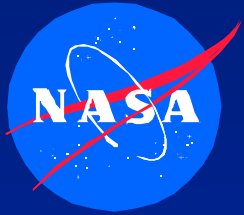




Early Warning: S-Curves

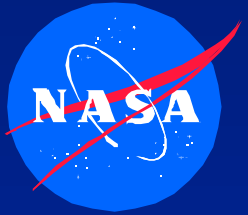
- ◆ Viewed with X-ray Imagery
- ◆ Based on Two Years of Observations
- ◆ Strong Correlation between Sigmoid Regions and CMEs
- ◆ Likely the Result of Twisted Solar Magnetic Fields
- ◆ May Provide Early Warning of Particle Events





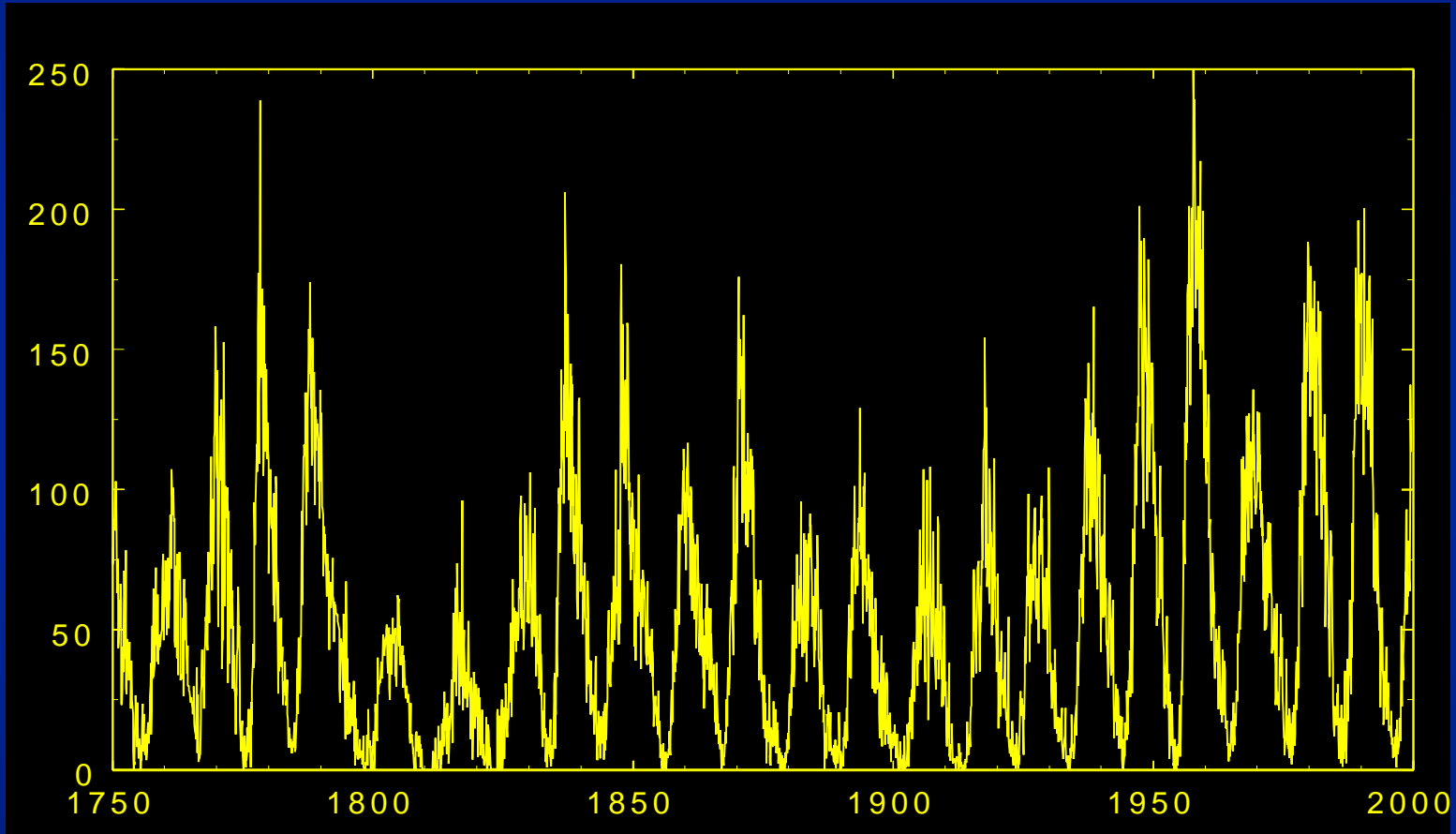
Solar Activity: Cyclic Variation

- ◆ Sunspot Cycle Discovered in mid 1800s
- ◆ Used as Indicator of Solar Activity
- ◆ Increased Solar Activity Results in:
 - » Increased Rate of CMEs
 - » Increased Rate of Solar Flares
 - » Increased Rate of Magnetic Storms
 - » Increased Levels of Electrons in Van Allen Belts
 - » Decreased Levels of Protons in Van Allen Belts
 - » Increased Incidence of New Belt Formation
 - » Decreased Levels of Galactic Cosmic Rays
 - » Increased Rate of Solar Particle Events
- ◆ Effects on Climate?

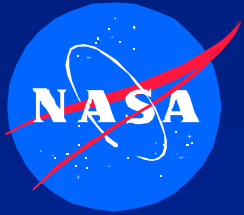


Sunspot Cycle

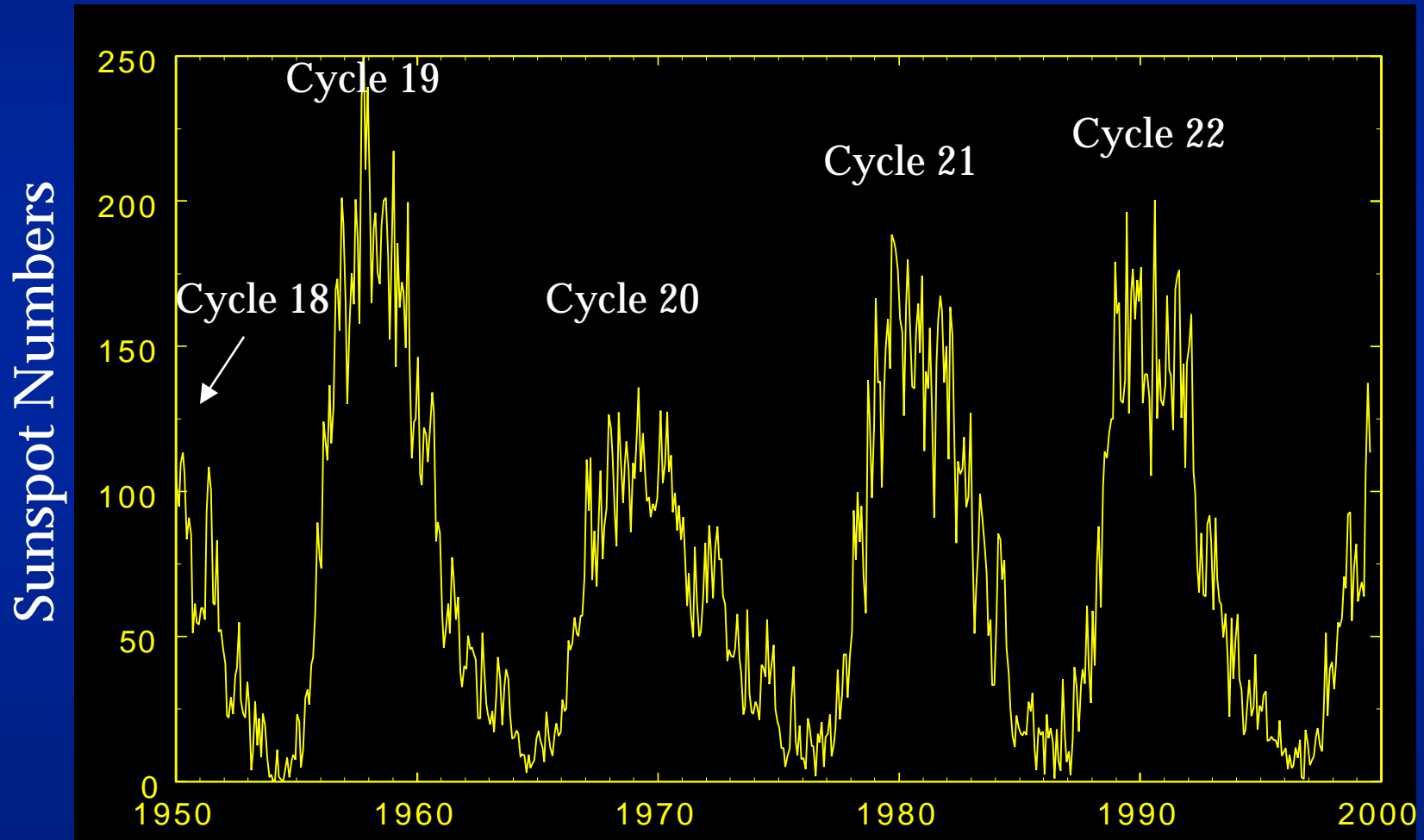
Sunspot Numbers

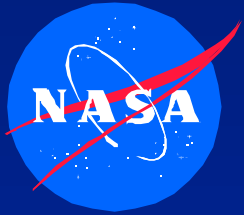


Length Varies from 9 - 13 Years
7 Years Solar Maximum, 4 Years Solar Minimum



Sunspot Cycle

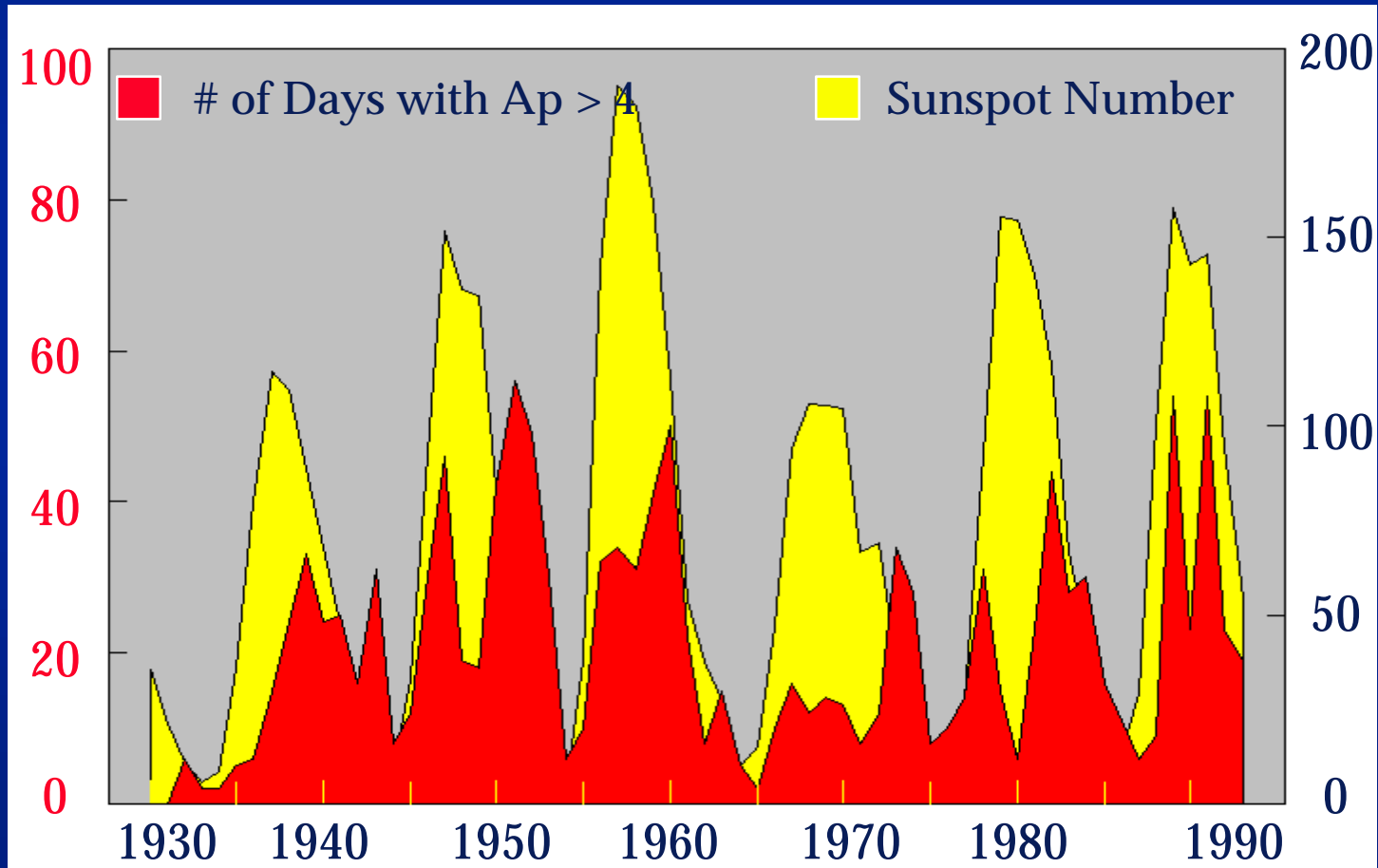




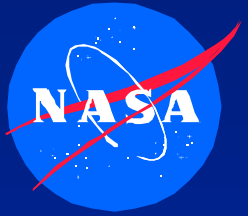
Sunspot Cycle with Magnetic Storms

Sunspots & Magnetic Storm Days

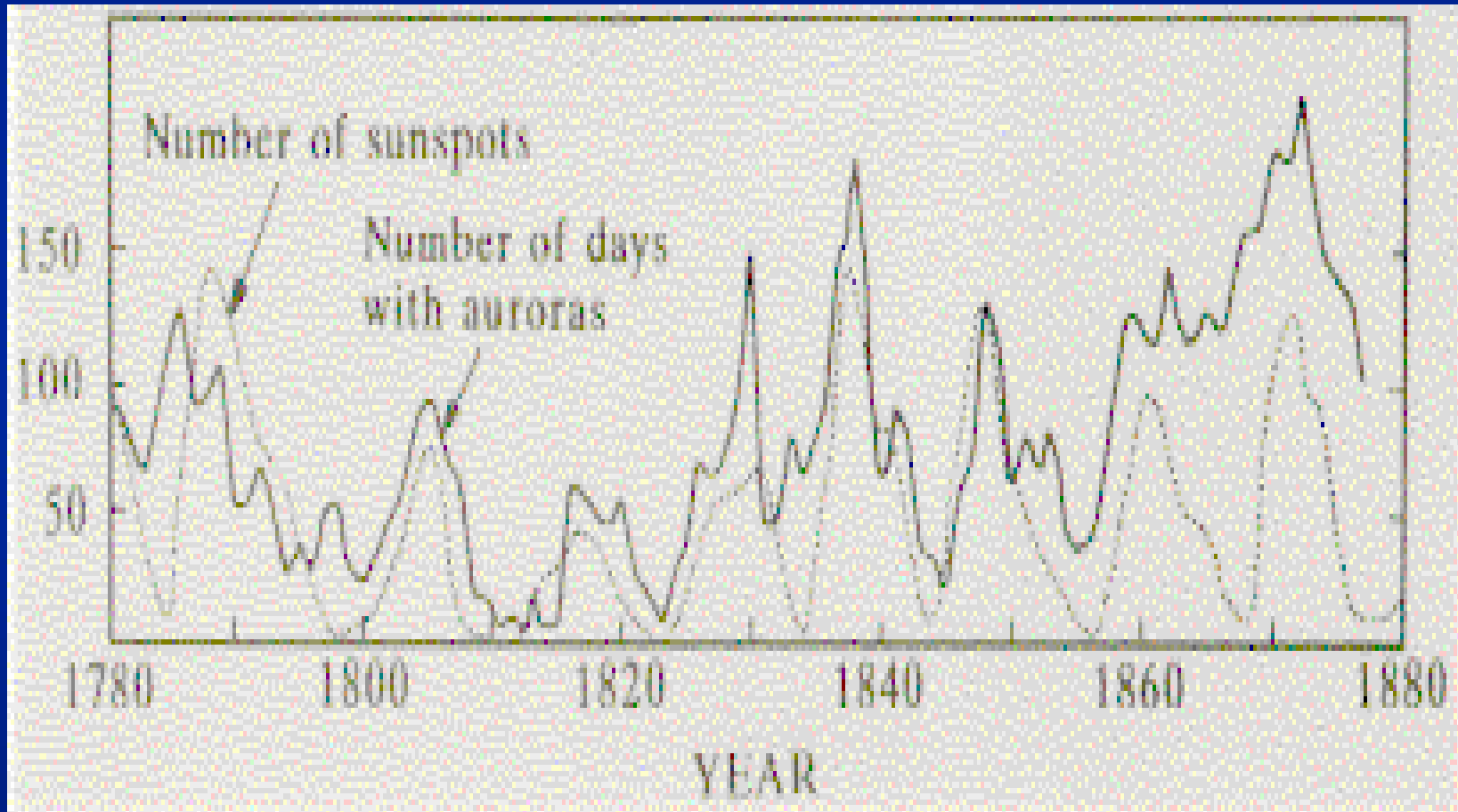
Annual Number of Days with $A_p > 4$



Annual Sunspot Number

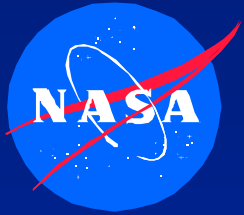


Solar Cycle - Aurora Days



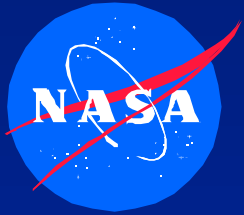
From Way North Magazine

J. Barth/Code 562



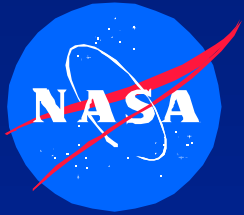
Discovery of Galactic Cosmic Rays - 1913

- ◆ **Electroscope Experiments**
 - » Dissipation of Charge on Leaves?
 - » Emissions from Materials on Earth
 - » “Clean” Instruments Did Not Eliminate Dissipation
- ◆ **Hess**
 - » Balloon Experiments with Electroscopes
 - » Hypothesis: Background Radiation Will Disappear with Increasing Altitude
 - » > 10,000 feet - Background Increased with Altitude
 - » Named “Cosmic Rays” by Hess



Galactic Cosmic Ray Ions

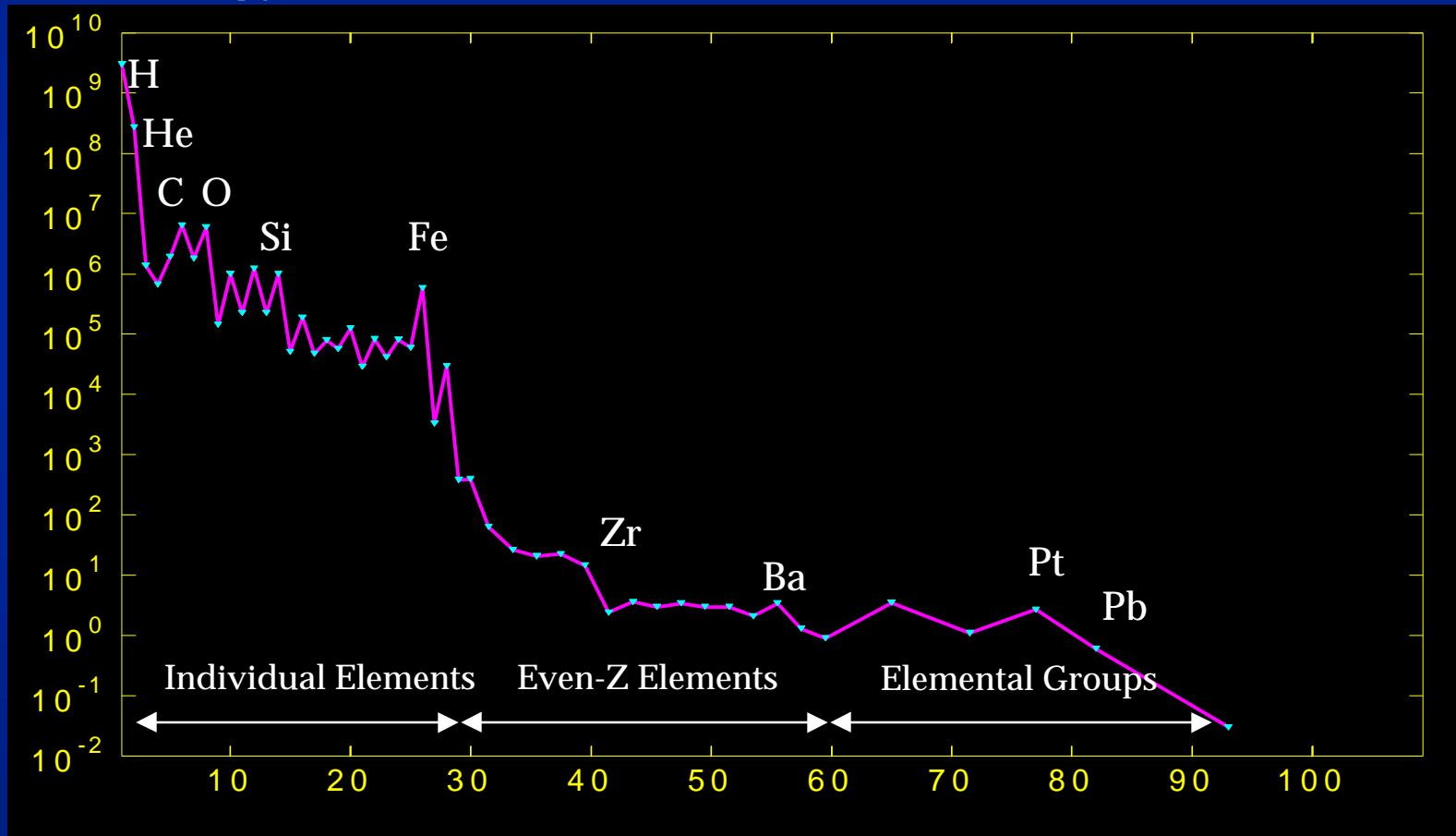
- ◆ All Elements in Periodic Table - 200 Million Years Old
- ◆ Energies in GeV
- ◆ Found Everywhere in Interplanetary Space
- ◆ Omnidirectional
- ◆ Mostly Fully Ionized - Protons & Bare Nuclei of Heavier Elements
- ◆ Cyclic Variation in Fluence Levels
 - » Lowest Levels = Solar Maximum Peak
 - » Highest Levels = Lowest Point in Solar Minimum
- ◆ Trajectories Bent by Magnetic Field
- ◆ Single Event Effects Hazard
- ◆ Model: CREME96



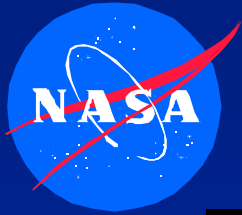
GCRs: Nuclear Composition

Energy = 2 GeV/n, Normalized to Silicon = 10^6

Relative Flux (Si = 10^6)



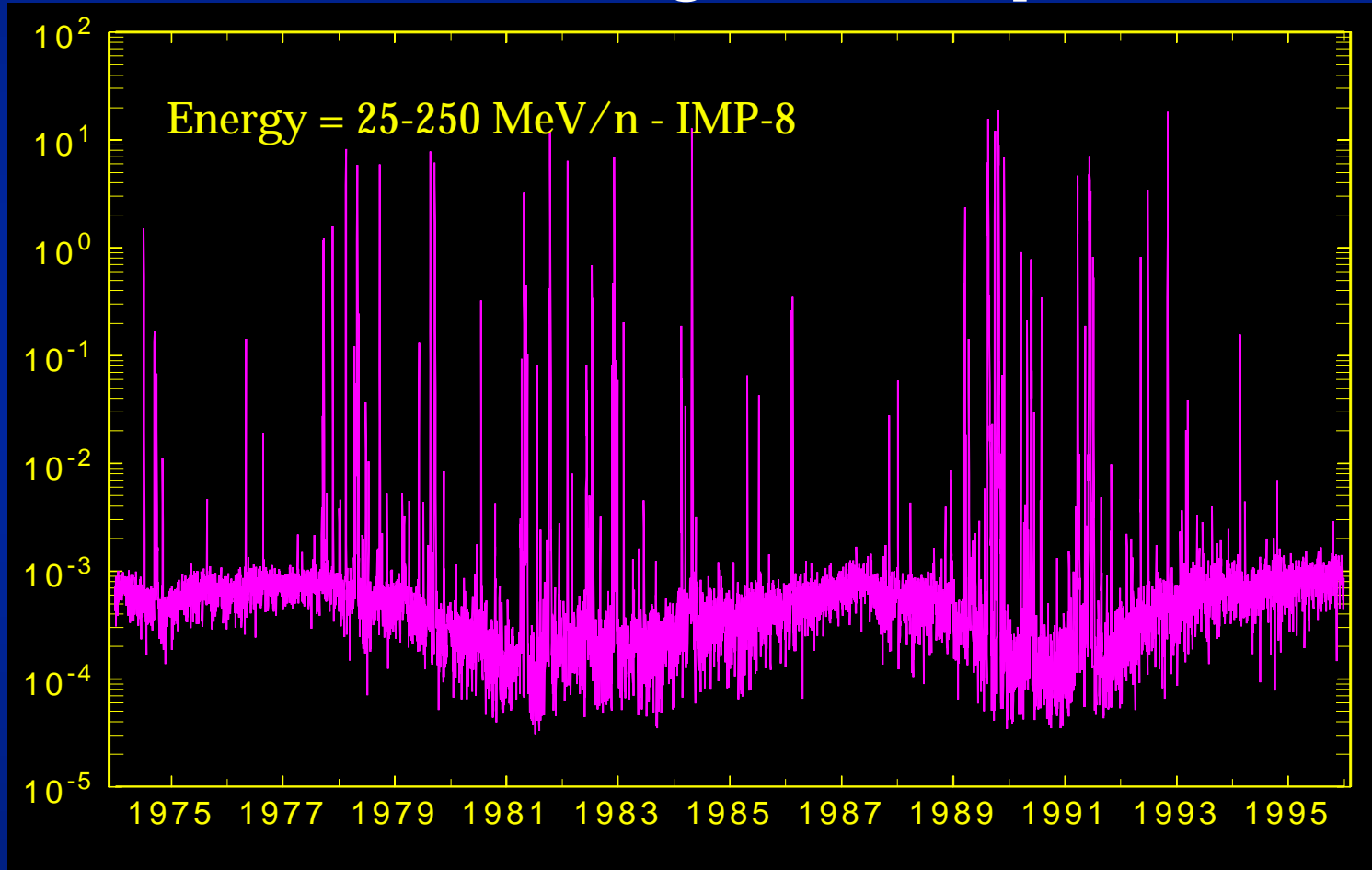
Nuclear Charge (Z)



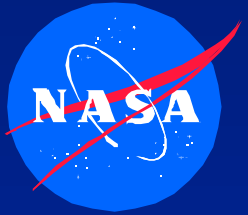
GCRs: Solar Modulation

CNO - 24 Hour Averaged Mean Exposure Flux

CNO (#/cm²/ster/s/MeV/n)



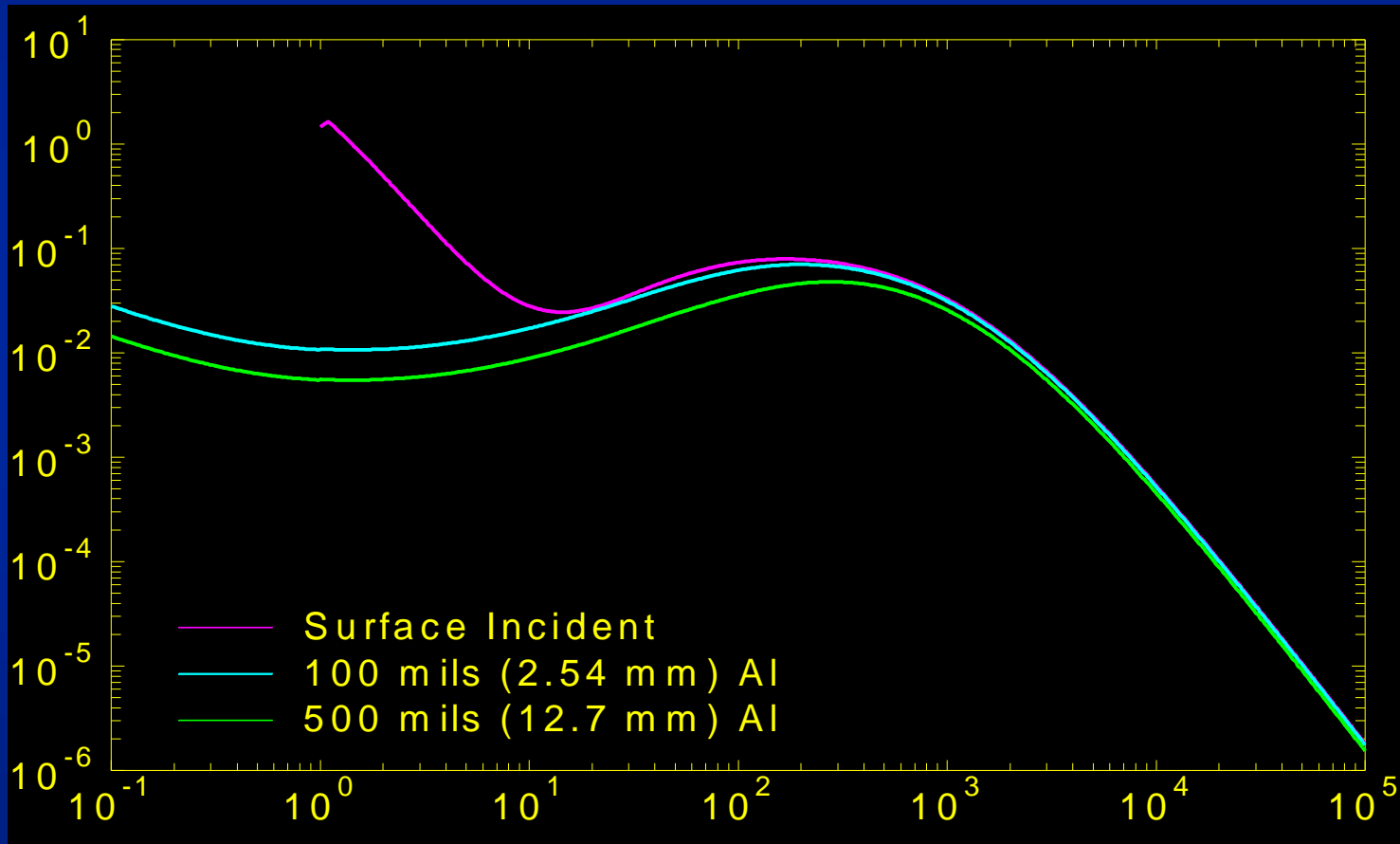
Date



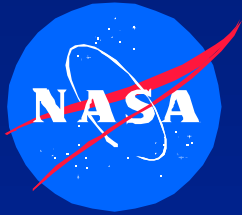
GCRs: Shielded Fluences - Fe

Interplanetary, CREME 96, Solar Minimum

Particles ($\# / \text{cm}^2 / \text{day} / \text{MeV} / \text{n}$)



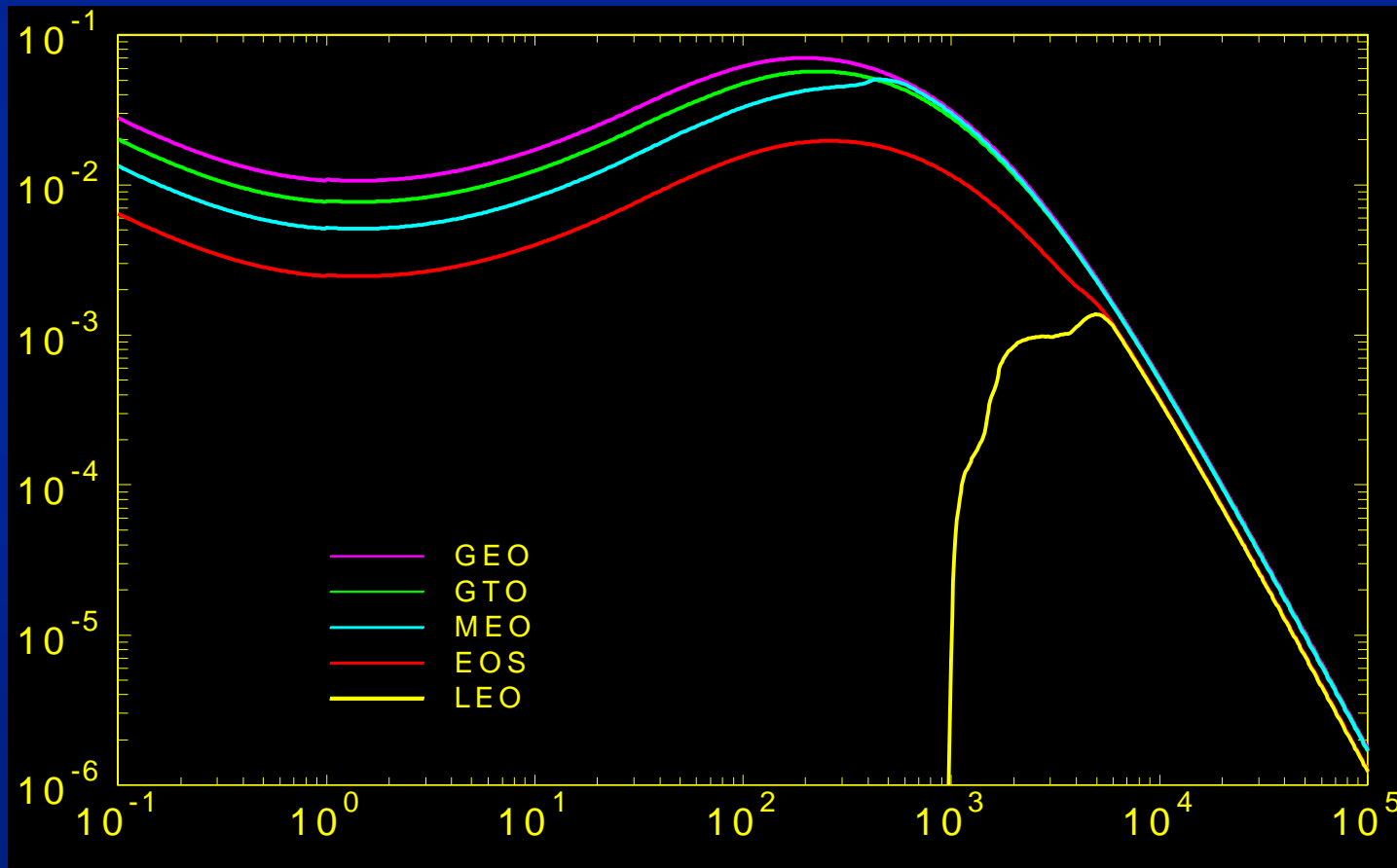
Energy (MeV/n)



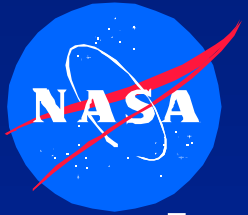
GCRs: Shielded Fluences - Fe

CREME 96, Solar Minimum, 100 mils (2.54 mm) Al

Particles ($\#/\text{cm}^2/\text{day}/\text{MeV}/\text{nuc}$)

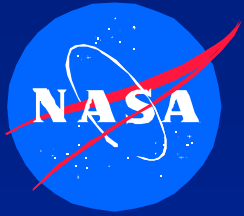


Energy (MeV/nuc)



Solar Particle Events

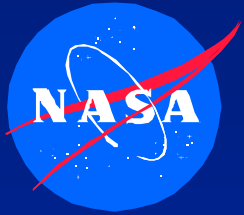
- ◆ Increased Levels of Protons & Heavier Ions
- ◆ Energies
 - » Protons - 100s of MeV
 - » Heavier Ions - 100s of GeV
- ◆ Abundances Dependent on Radial Distance from Sun
- ◆ Partially Ionized - Greater Ability to Penetrate Magnetosphere
- ◆ Number & Intensity of Events Increases Dramatically During Solar Maximum
- ◆ Models
 - » Dose - SOLPRO, JPL, Xapsos/NRL
 - » Single Event Effects - CREME96 (Protons & Heavier Ions)



Solar Particle Events

“The Solar Flare Myth” - Gosling

- ◆ Poor Correlation with Solar Flares
- ◆ Strong Correlation with Coronal Mass Ejections
 - » No Fundamental Association with Flares
 - » Transient Shock Wave Disturbances in the Solar Wind
 - » Large Geomagnetic Storms
 - » Large Particle Events

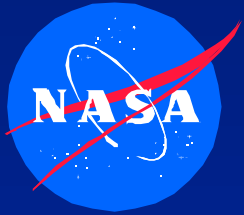


Solar Particle Events

“A New Paradigm” - Reames

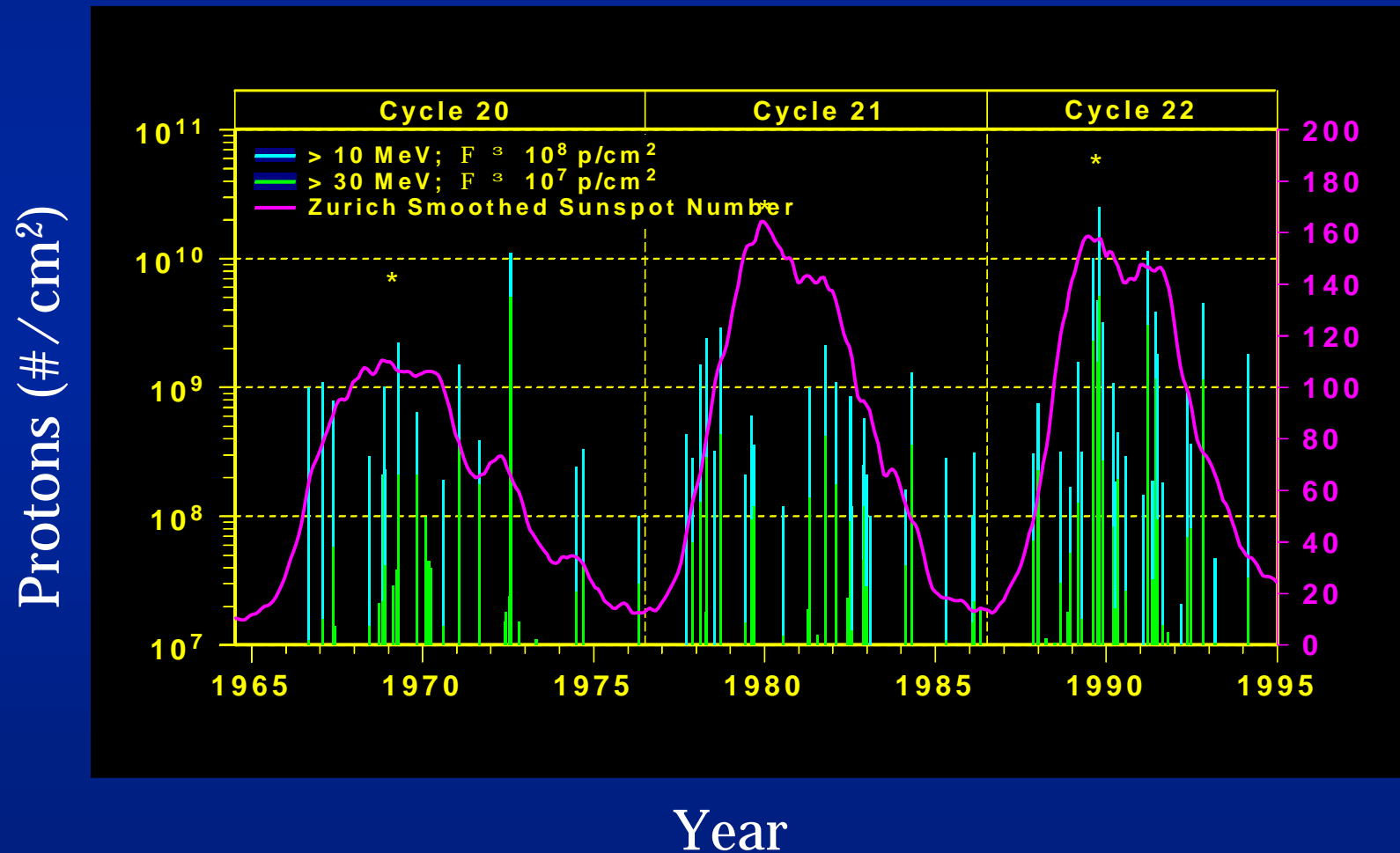
2 Types of Events

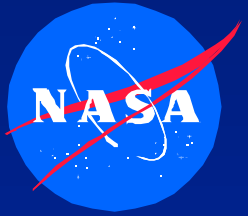
- ◆ Gradual ~ 10 per Year
 - » Coronal Mass Ejection Driven Shocks
 - » Same Elemental Abundances & Ionization States as Coronal & Solar Wind Plasma
- ◆ Impulsive ~ 1000 per Year
 - » Flare Accelerated
 - » Abundances Characteristic of Interactions in the Flare Plasma



Sunspot Cycle with Solar Proton Events

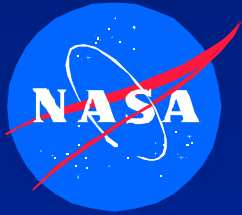
Proton Event Fluences





Modeling Approach

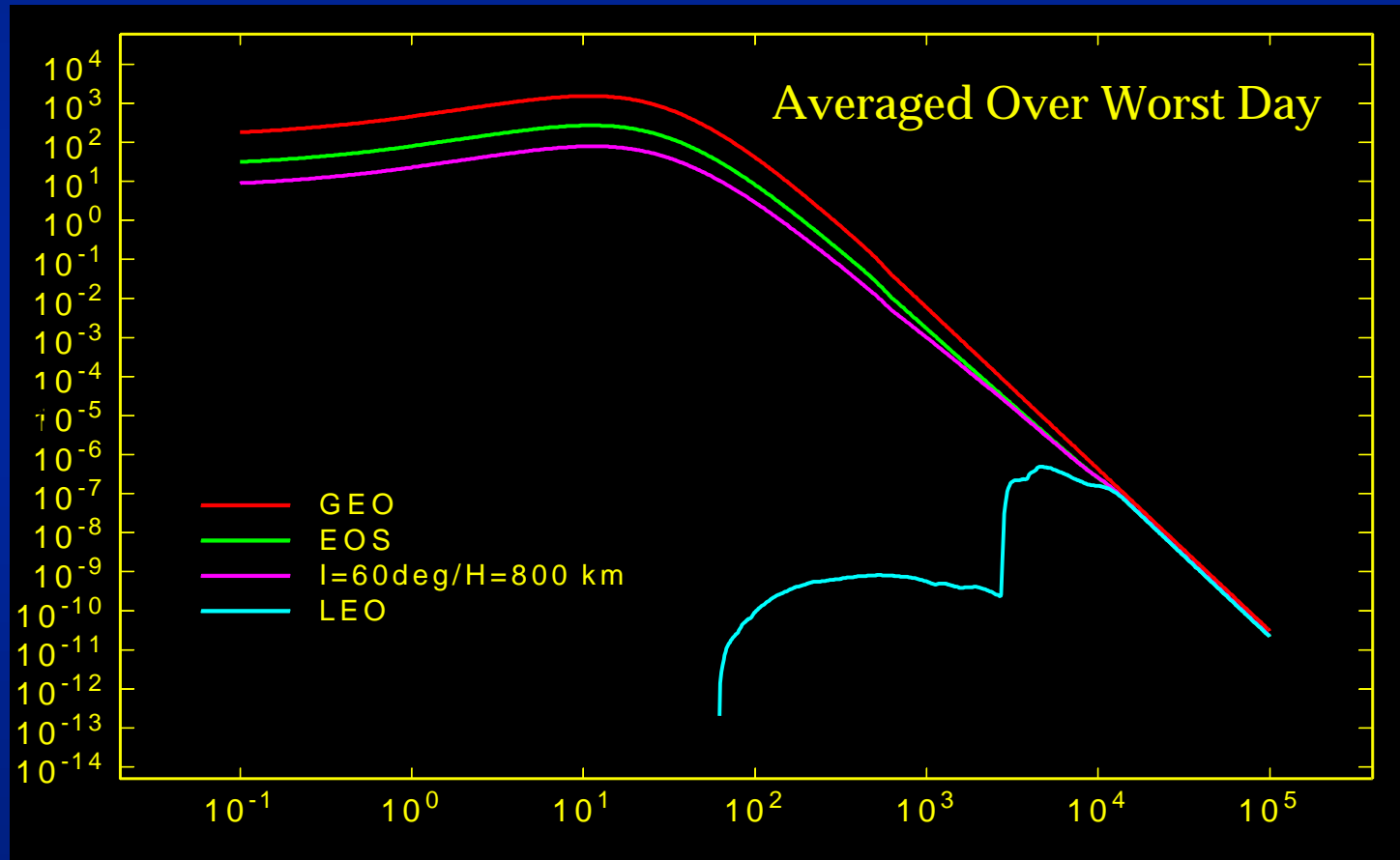
- ◆ Statistical Engineering Model
 - » Intensity as a Function of Mission Duration & Confidence Level
 - » Does Not Predict When Events Occur
- ◆ Use Maximum Entropy Principle - Incomplete Data Set
 - » Determines Frequency Distribution of Large Solar Proton Events
 - » Frequency Distribution Consistent with Other Complex Physical Phenomena such as Earthquakes
- ◆ Use Extreme Value Theory
 - » Determines Upper Limit for Occurrence of Huge Events
 - » New Upper Limit Consistent with Data Sets Dating Back to Ancient Times - Lunar Rock Record



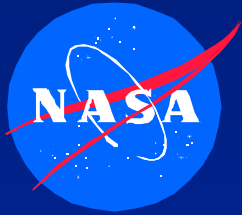
Solar Protons: Orbits

Proton Levels Predicted by CREME 96

Protons (# / cm² / sec / MeV)



Energy (MeV)



TIROS Measurement of Protons

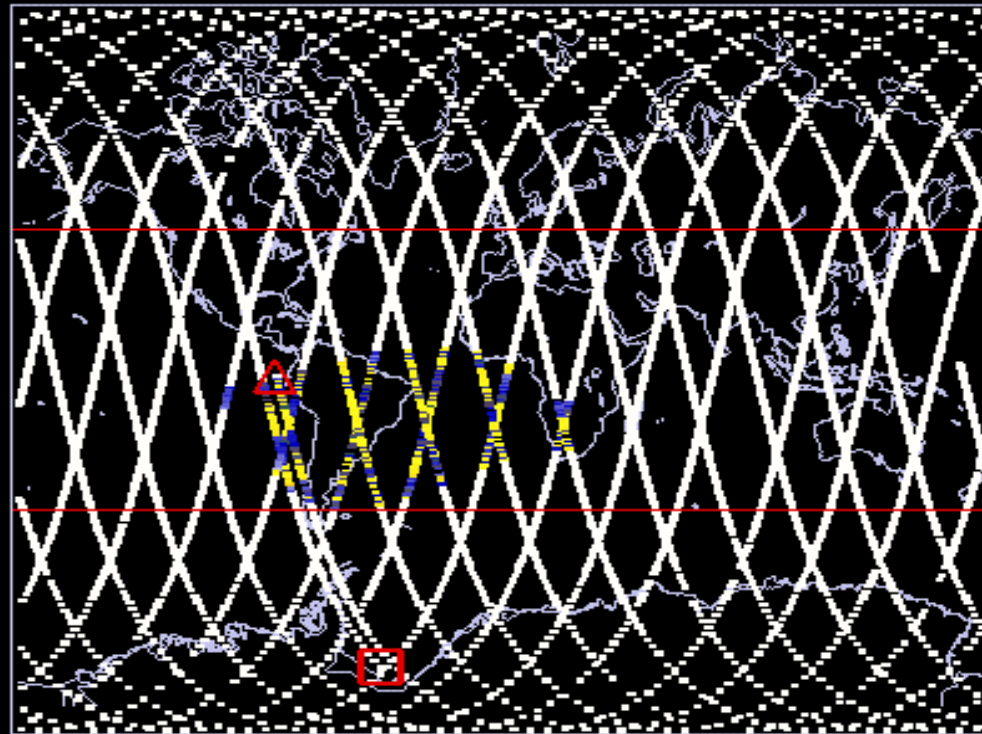
Day Before Coronal Mass Ejection

H= 870/870 km

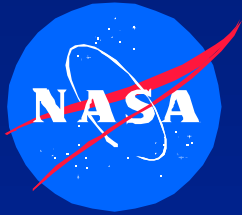
NOAA-12 SEM
5 November 1997

BELT INDICES:
Total 0.9892
Inner 0.9893
Slot {Insig}
Outer {Insig}

80-250 MeV Protons (Omnidirectional)



NOAA Space Environment Center



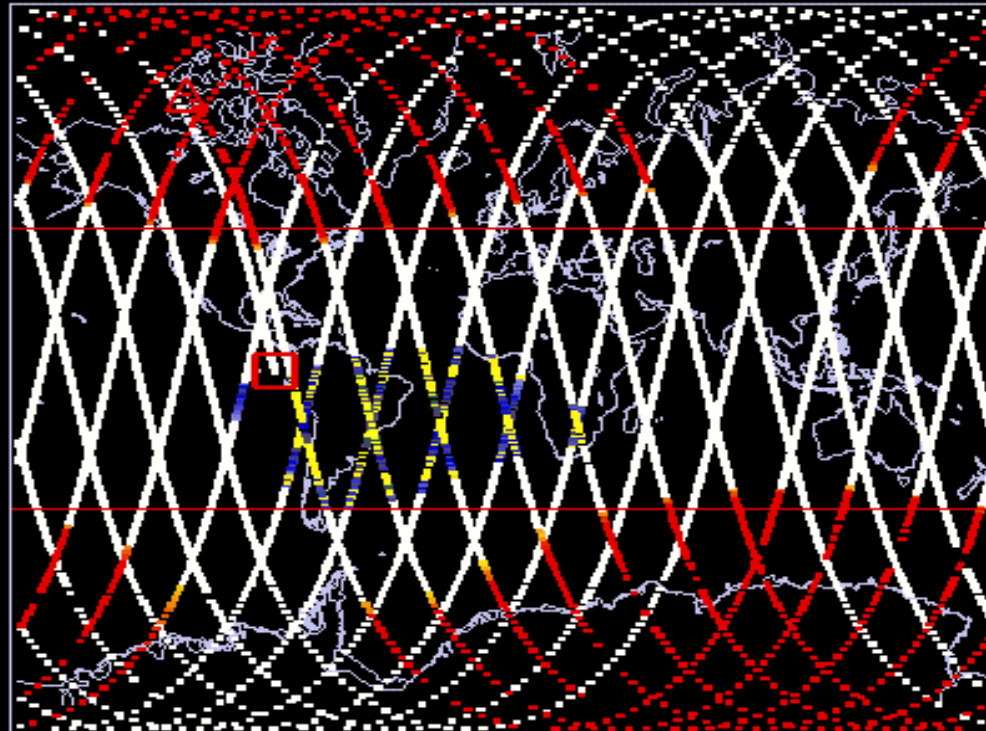
TIROS Measurement of Protons

November 6, 1997 Coronal Mass Ejection

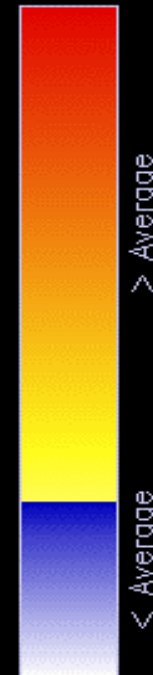
H= 870/870 km

NOAA-12 SEM
5 November 1997

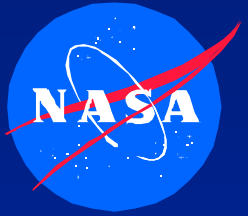
80-250 MeV Protons (Omnidirectional)



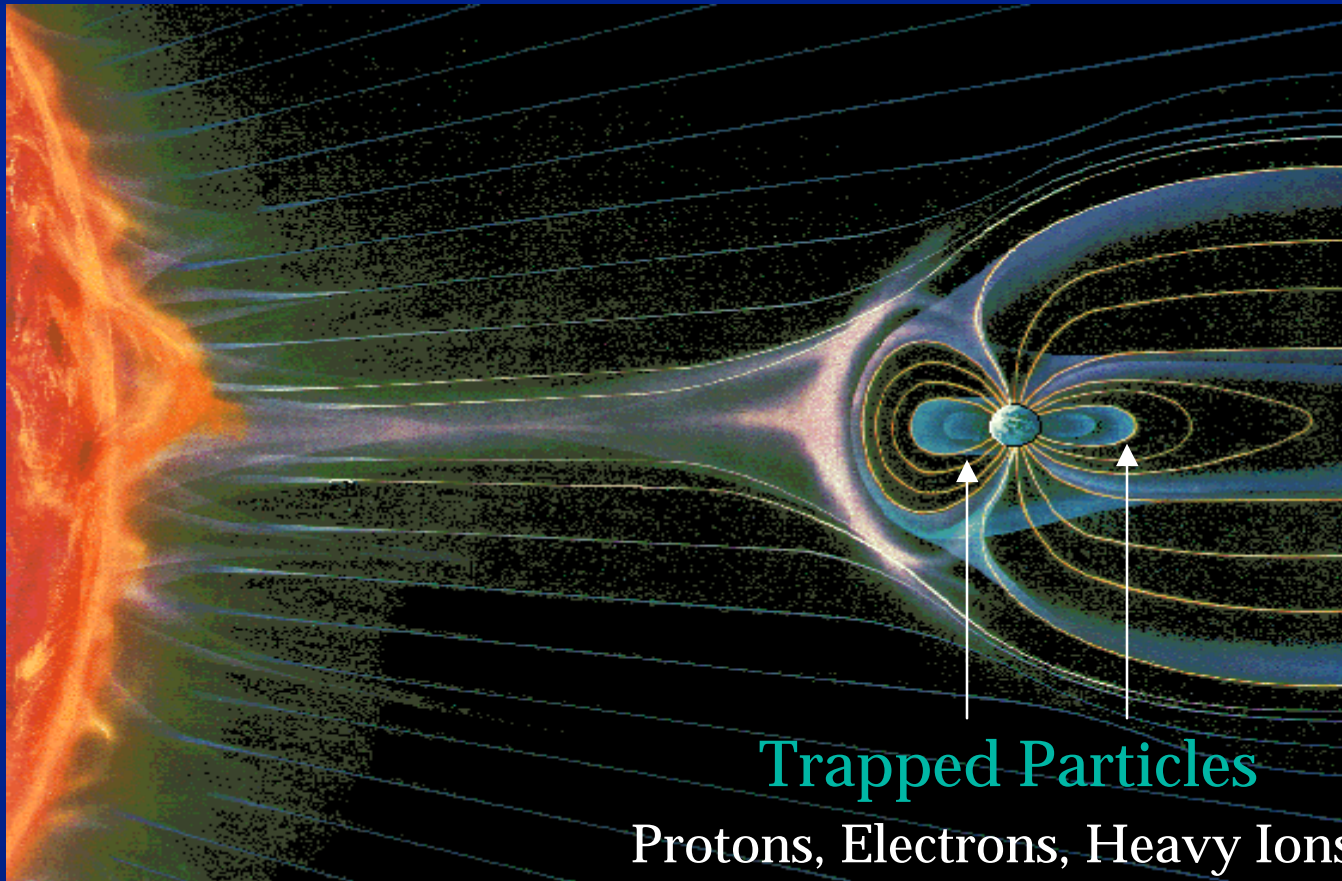
BELT INDICES:
Total 1.3799
Inner 0.9909
Slot (Insig)
Outer 13.3677



NOAA Space Environment Center

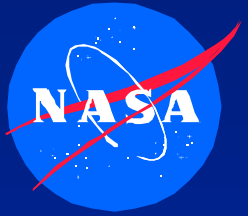


Trapped Radiation



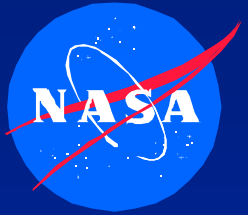
Nikkei Science, Inc. of Japan, by K. Endo

J. Barth/Code 562



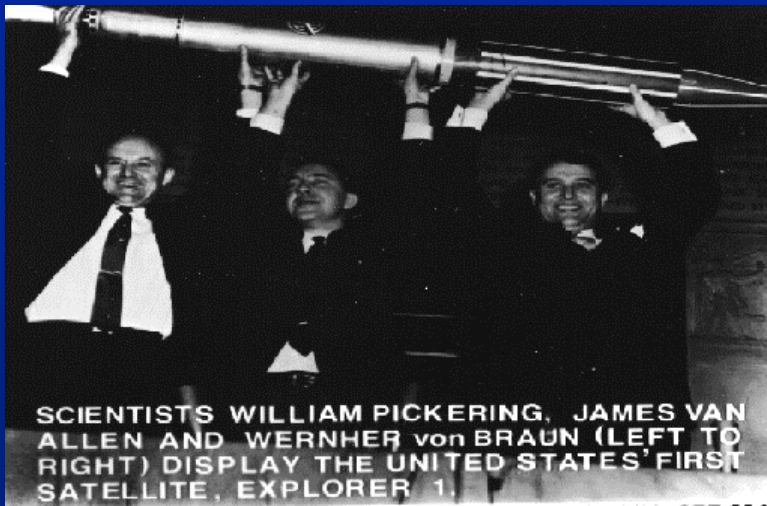
Charged Particle Motion

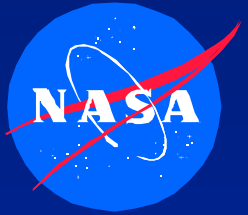
- ◆ Birkeland - 1895
 - » Vacuum chamber experiments to study aurora
 - » With Poincare showed that charged particles spiraled around field lines and are repelled by strong fields
- ◆ Störmer -
 - » Continued work of Birkeland on aurora
 - » Calculations led to theory that there was a belt-like area around the earth in which particles were reflected back and forth between the poles
- ◆ Singer (U. of Md) - 1957
 - » Proposed that ring current could be carried by lower energy particles injected by into trapped orbits by magnetic storms
- ◆ Christofilos
 - » Study of particle motion in magnetic fields - Project Argus



Discovery of the Radiation Belts

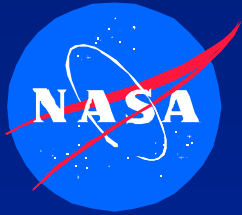
- ◆ James Van Allen
 - » First Observation of Auroral Electrons with a Rocket
 - » Cosmic Ray Detector
- ◆ Highlight of US Participation in IGY





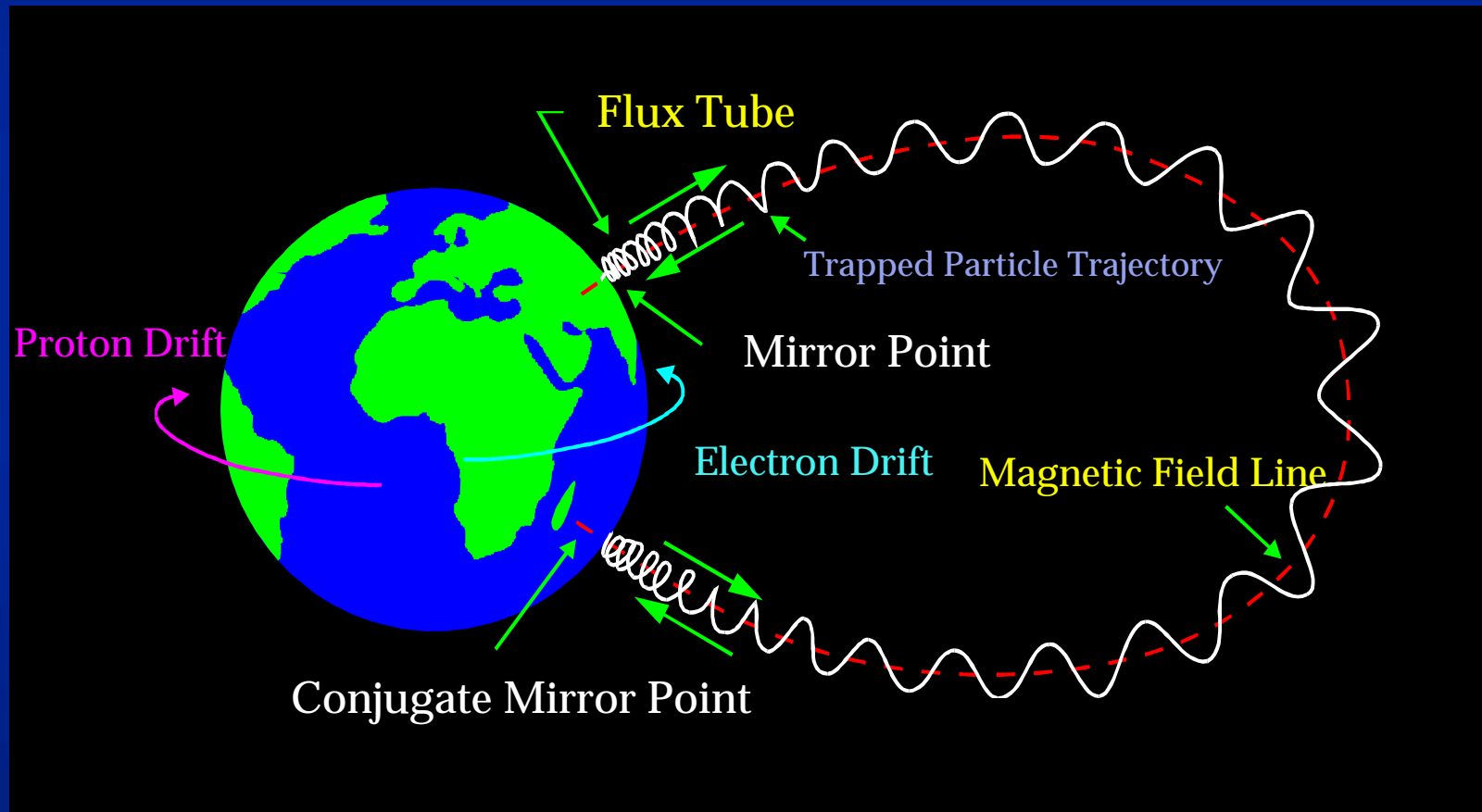
Trapped - Van Allen Belts

- ◆ Omnidirectional
- ◆ Components
 - » Protons: $E \sim .04 - 500 \text{ MeV}$
 - » Electrons: $E \sim .04 - 7(?) \text{ MeV}$
 - » Heavier Ions: Low E - Non-problem for Electronics
- ◆ Location of Peak Levels Depends on Energy
- ◆ Average Counts Vary Slowly with the Solar Cycle
- ◆ Location of Populations Shifts with Time
- ◆ Counts Can Increase by Orders of Magnitude During Magnetic Storms
 - » March 1991 Storm - Increases Were Long Term

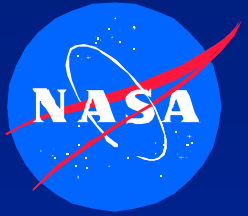


Trapped Particle Motions

Spiral, Bounce, Drift

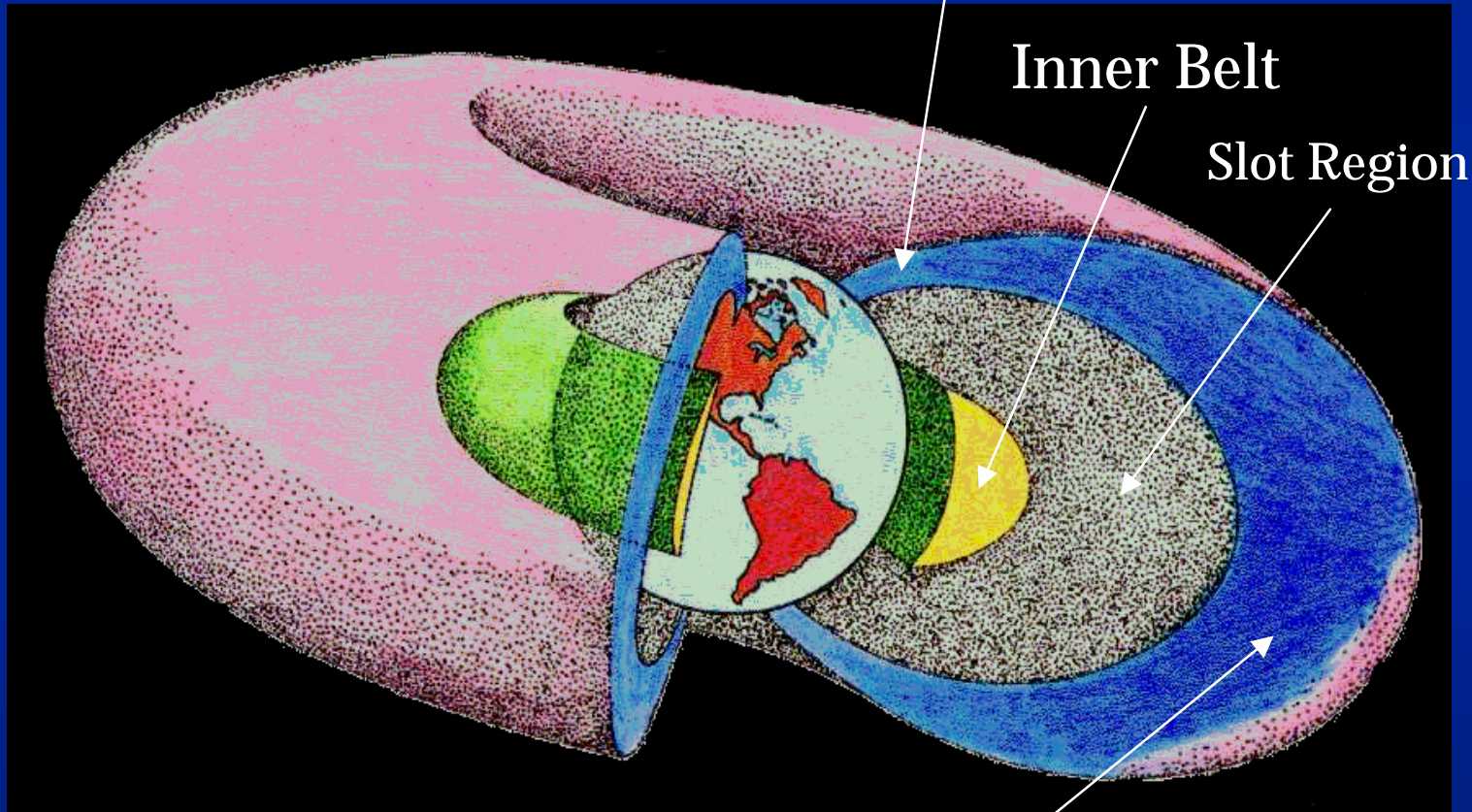


after Hess



Van Allen Belts

High Latitude Horns

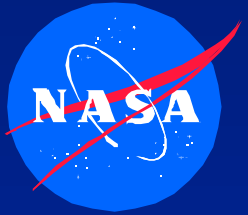


Inner Belt

Slot Region

Outer Belt

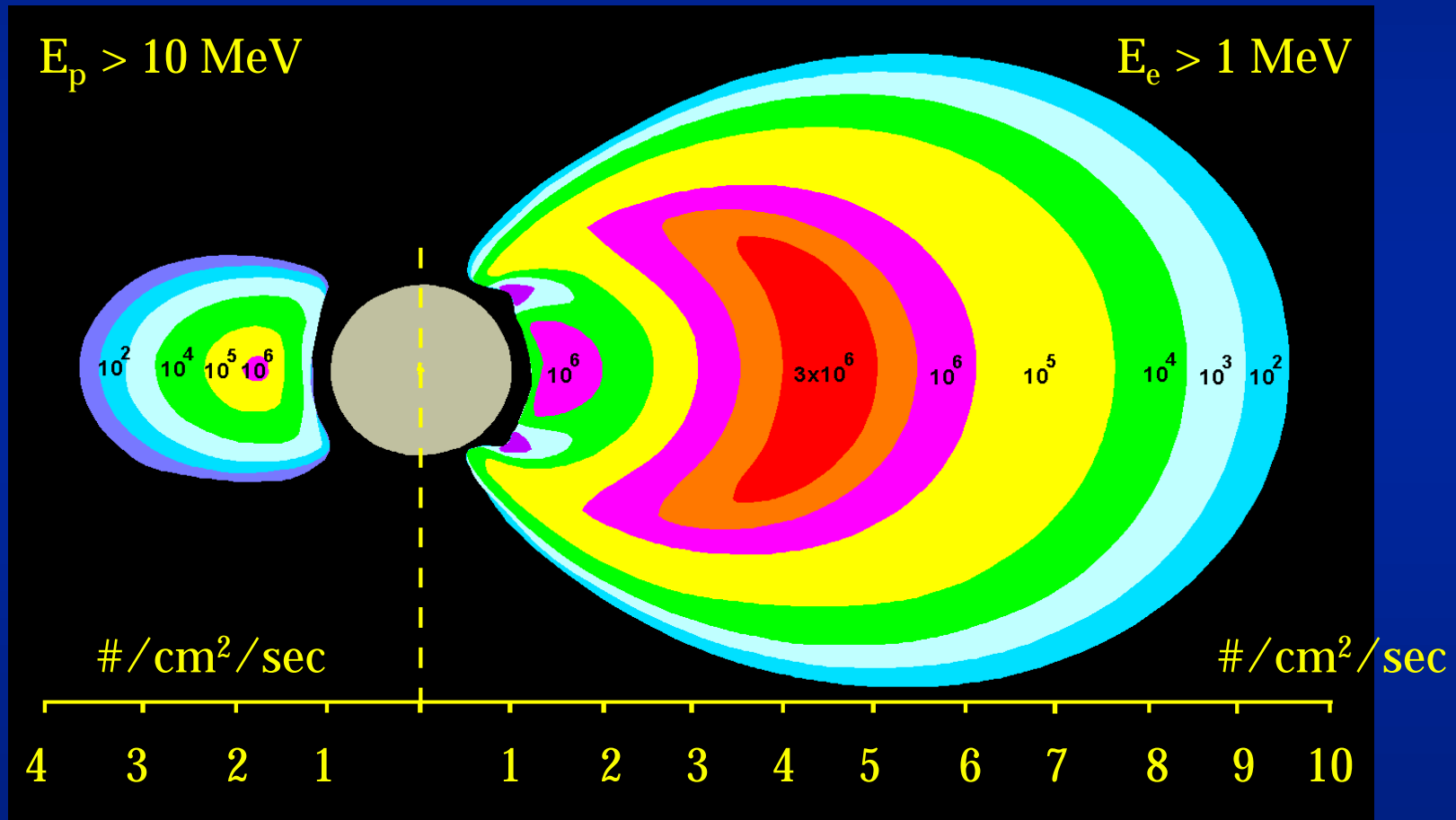
BIRA/IASB

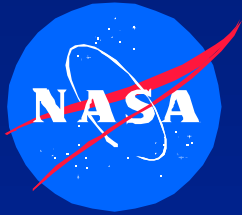


Proton & Electron Intensities

AP-8 Model

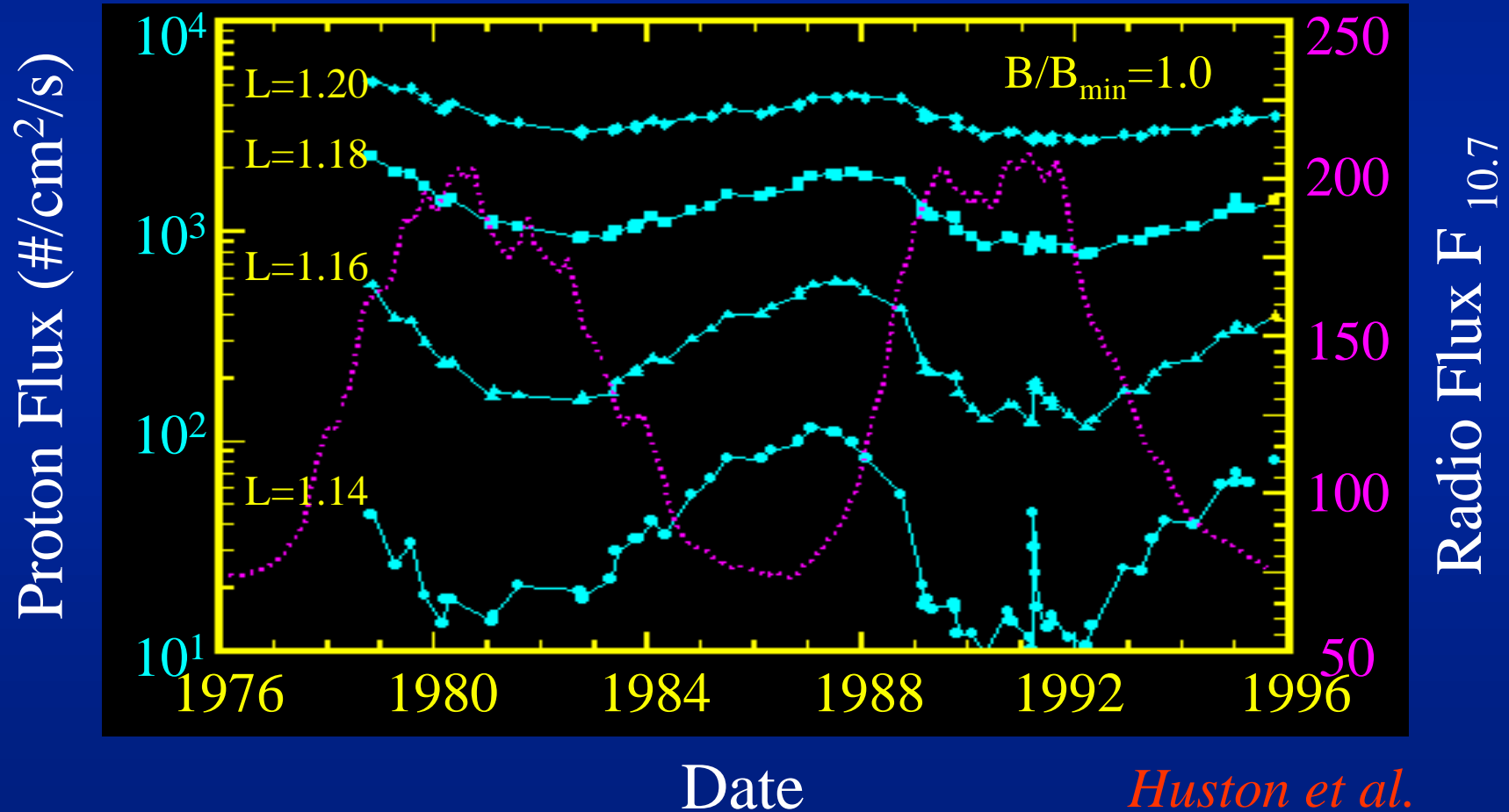
AE-8 Model

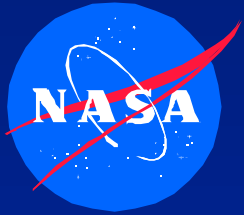




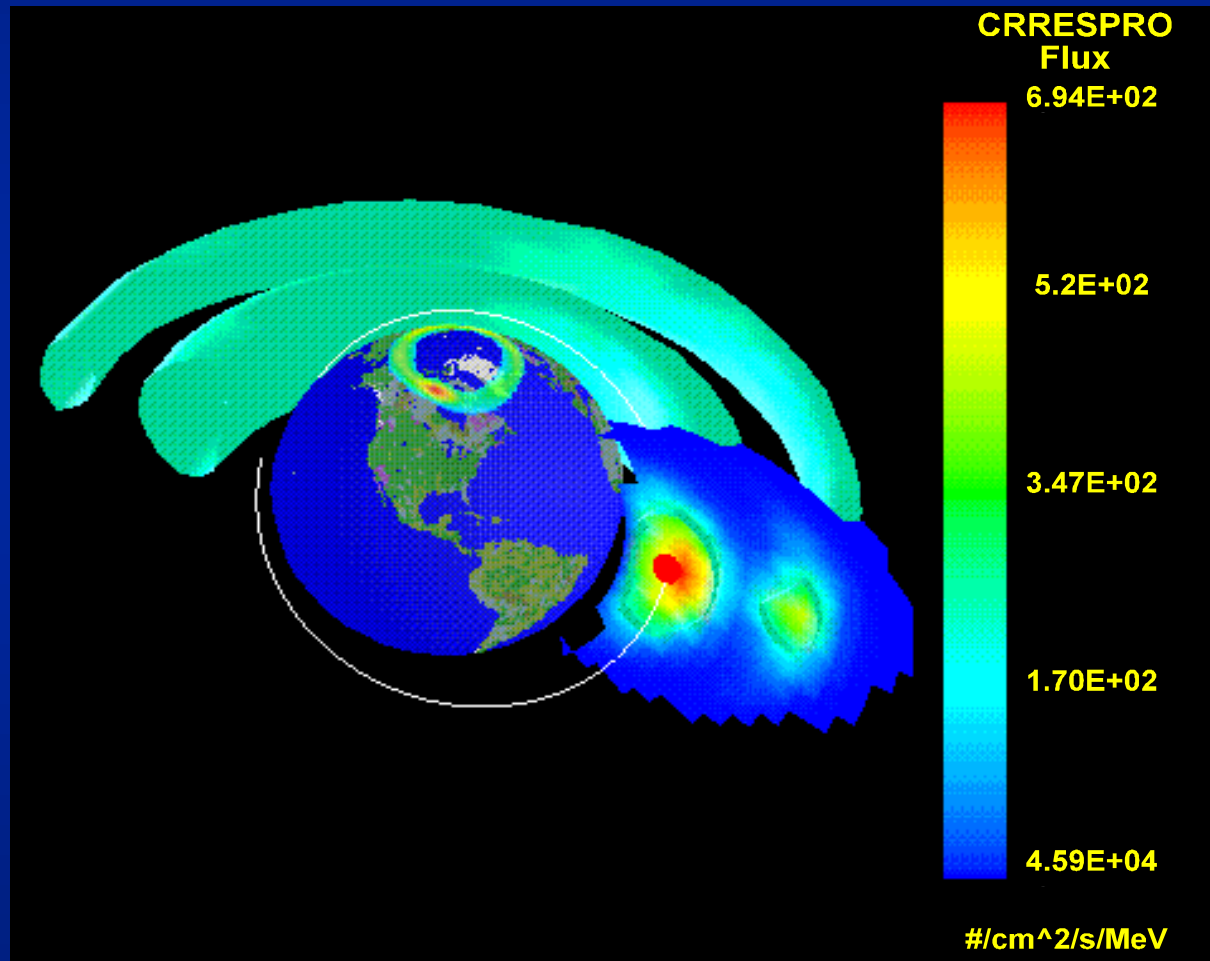
TIROS/NOAA Trapped Protons

Solar Cycle Variation: 80-215 MeV Protons



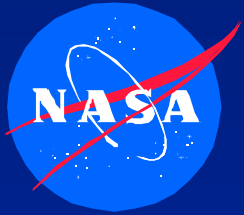


CRRES - Measured Proton Belt



J. Barth/Code 562

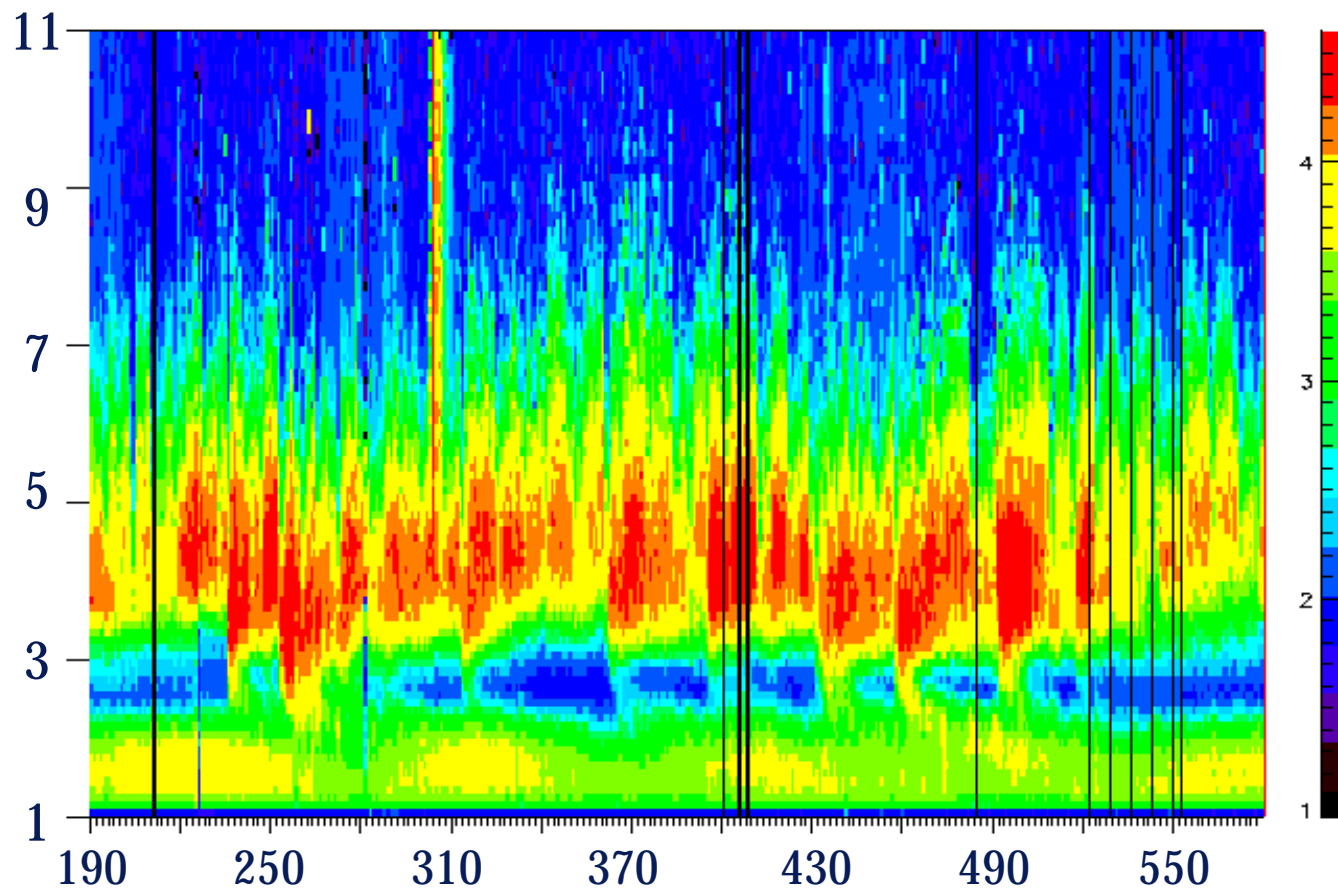
AF Phillips Laboratory, SPD/GD



Activity in the Slot Region - SAMPEX

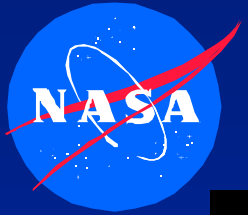
SAMPEX/P1ADC: Electrons $E > 0.4$ MeV

L-Shell



J. Barth/Code 562

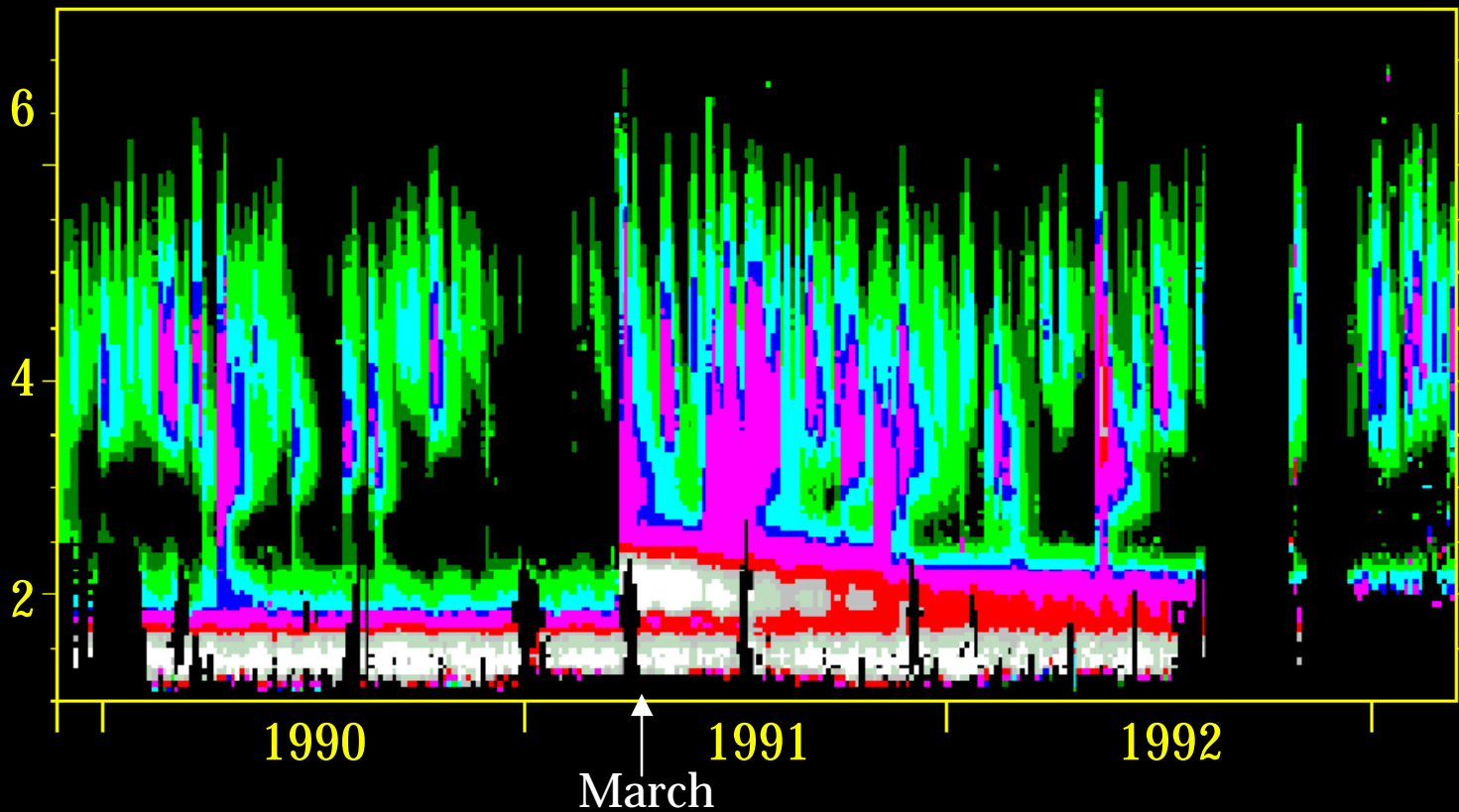
Day (1992)



Magnetic Storms - Hipparcos

Star Mapper - Radiation Background

L-Shell



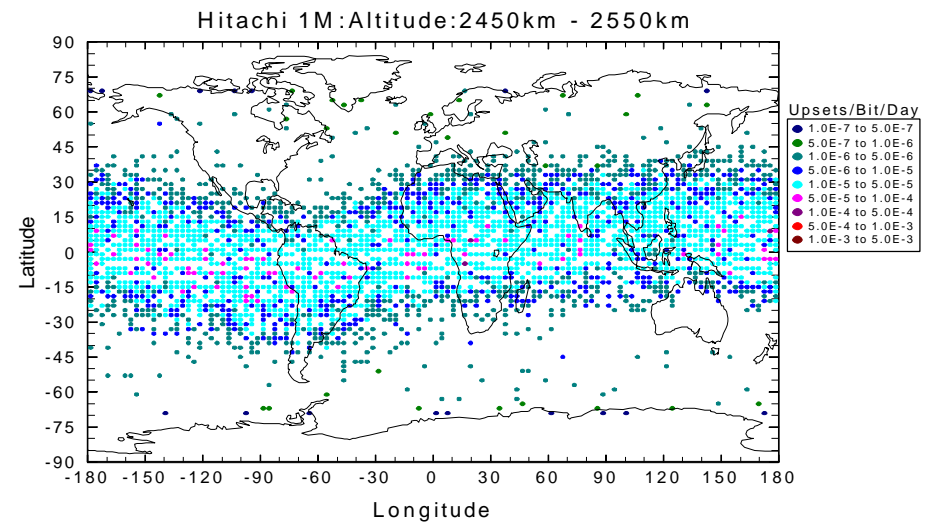
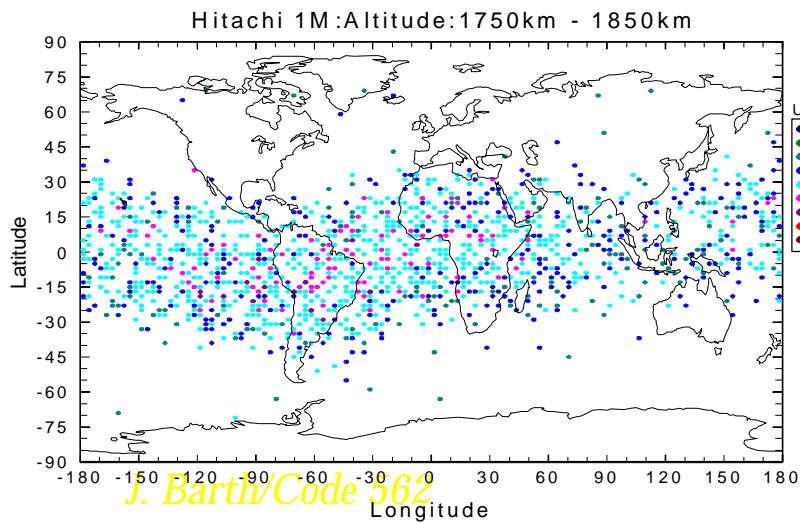
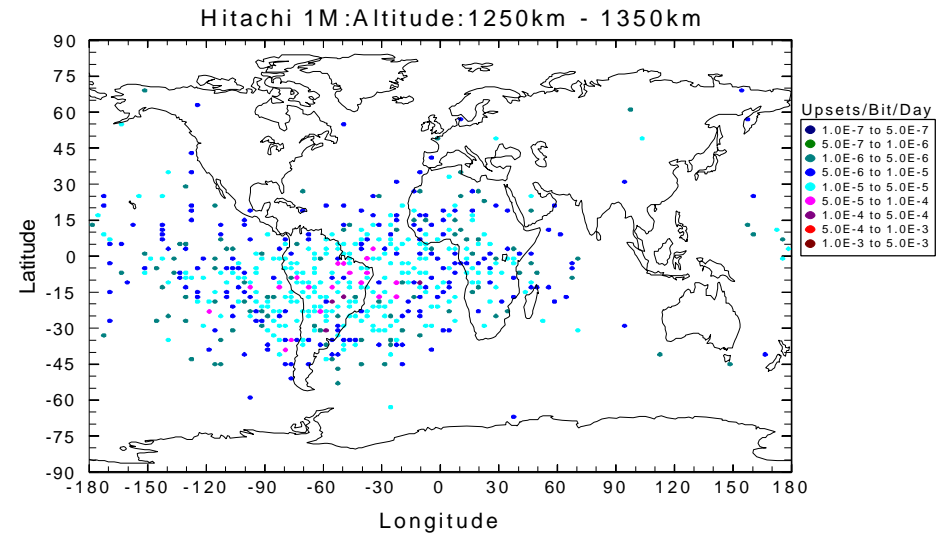
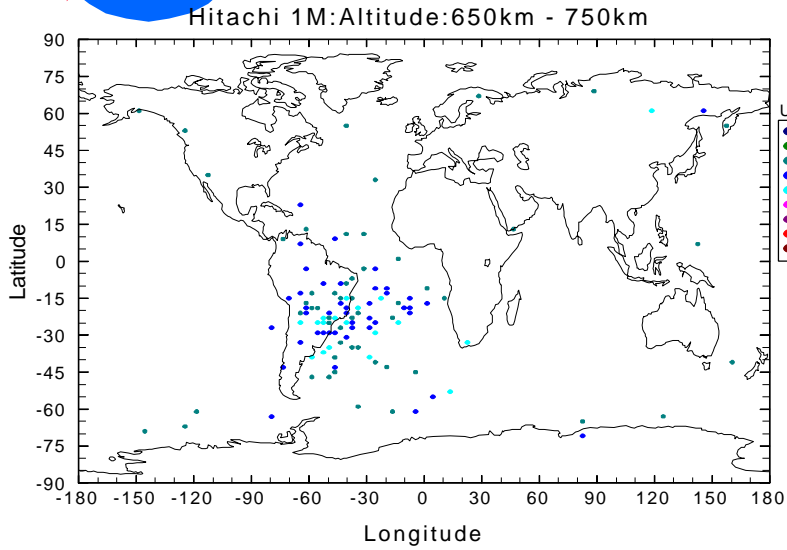
4-Day, 9-Orbit Averages

Daly, et al.

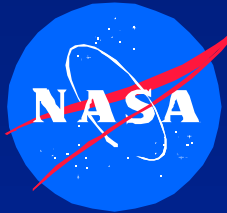
J. Barth/Code 562



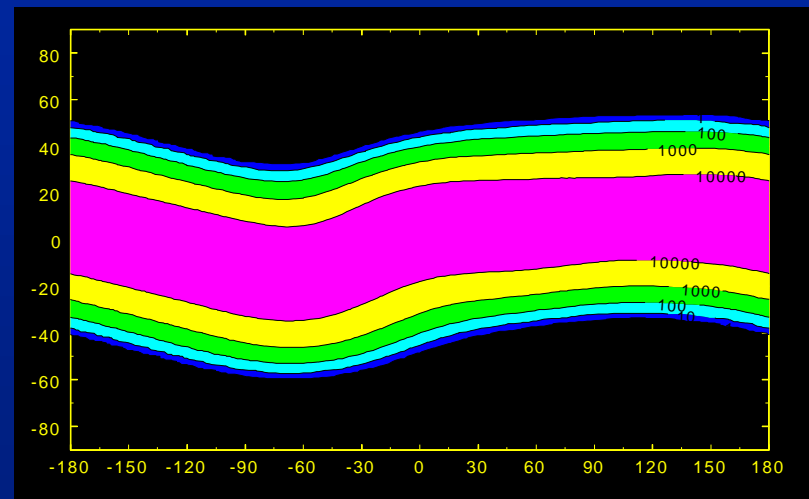
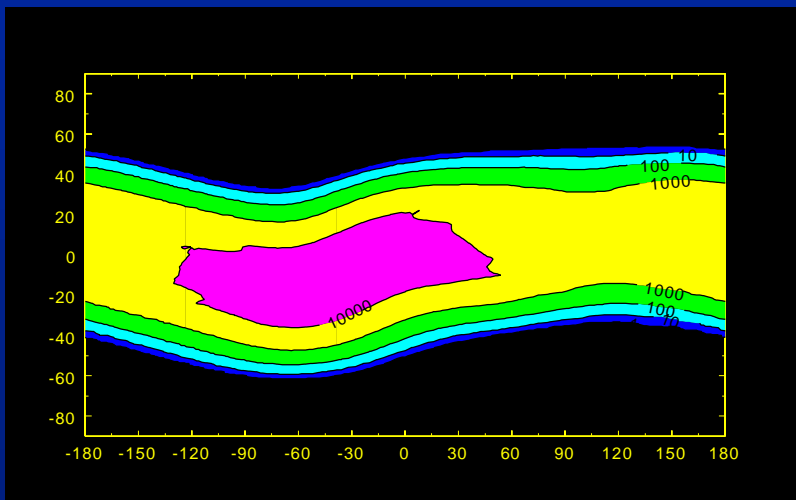
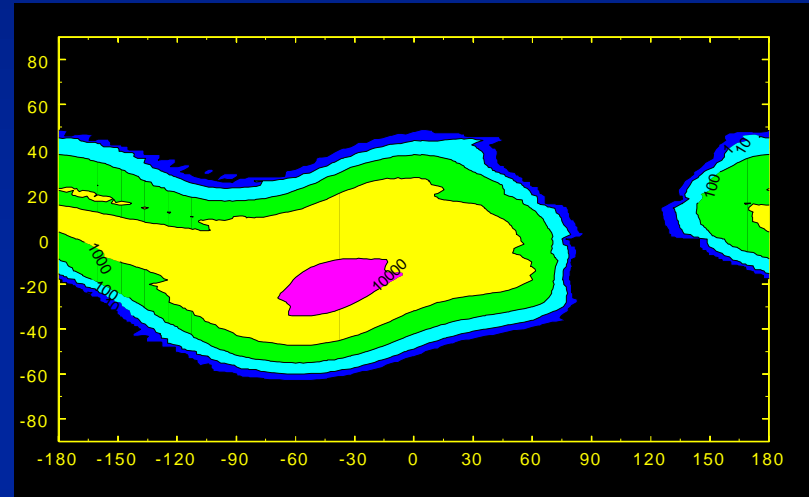
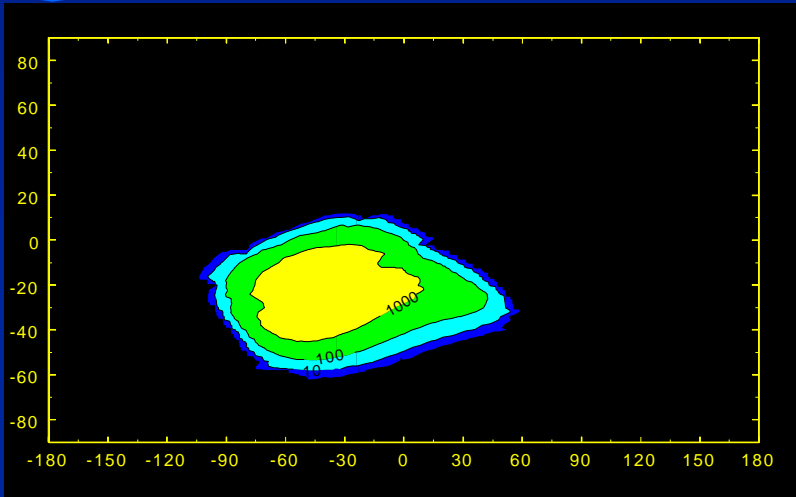
SRAM Upset Rates on CRUX/APEX

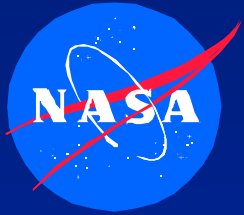


J. Barth/Code 362



Trapped Protons for $E > 30$ MeV (#/cm²/s) - Solar Minimum

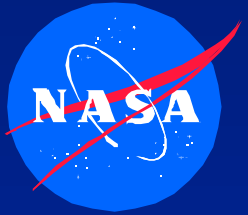




Particle Cascades in Atmosphere

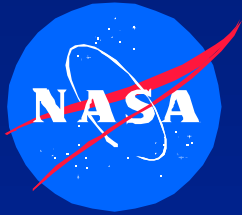
- ◆ Collisions between cosmic rays & atmospheric O & N
- ◆ Important product is neutrons
 - » Single Event Upsets
 - Shuttle
 - Aircraft
 - Ground
 - » Passenger & crew exposure in aircraft





Neutrons

- ◆ Source - Secondary Products of Particle Cascades
 - » Spacecraft Materials
 - » Galactic Cosmic Ray Collisions with Atmospheric O & N
- ◆ Single Event Upset Hazard
 - » Ground Level in Large Memory Banks
 - » Avionics
 - » Low Earth Orbits - Shuttle
- ◆ First Recognized as Problem in 1980s



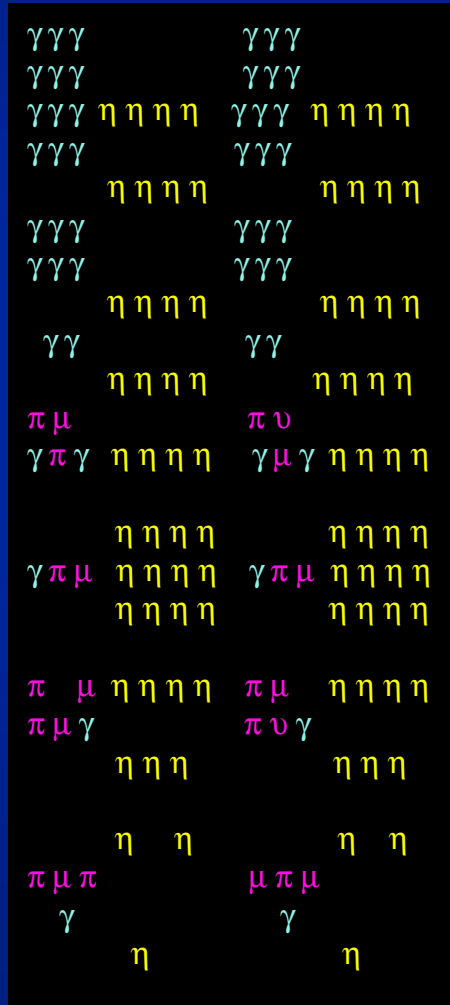
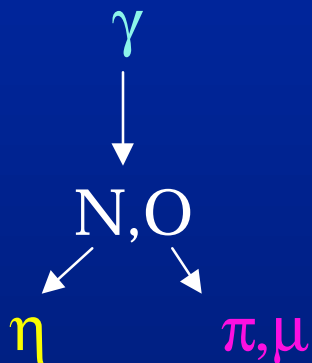
Neutron Environment

Normand et al.

γ Primary Cosmic Rays

η Neutrons

π, μ Secondary Cosmic Rays



← 1,000,000 feet
330 km

Shuttle

← 150,000 feet
50,000 m

Top of Atmosphere

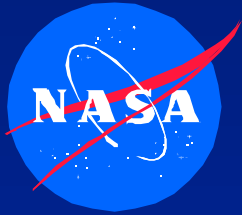
← 60,000 feet
20,000 m

Peak Neutron Flux

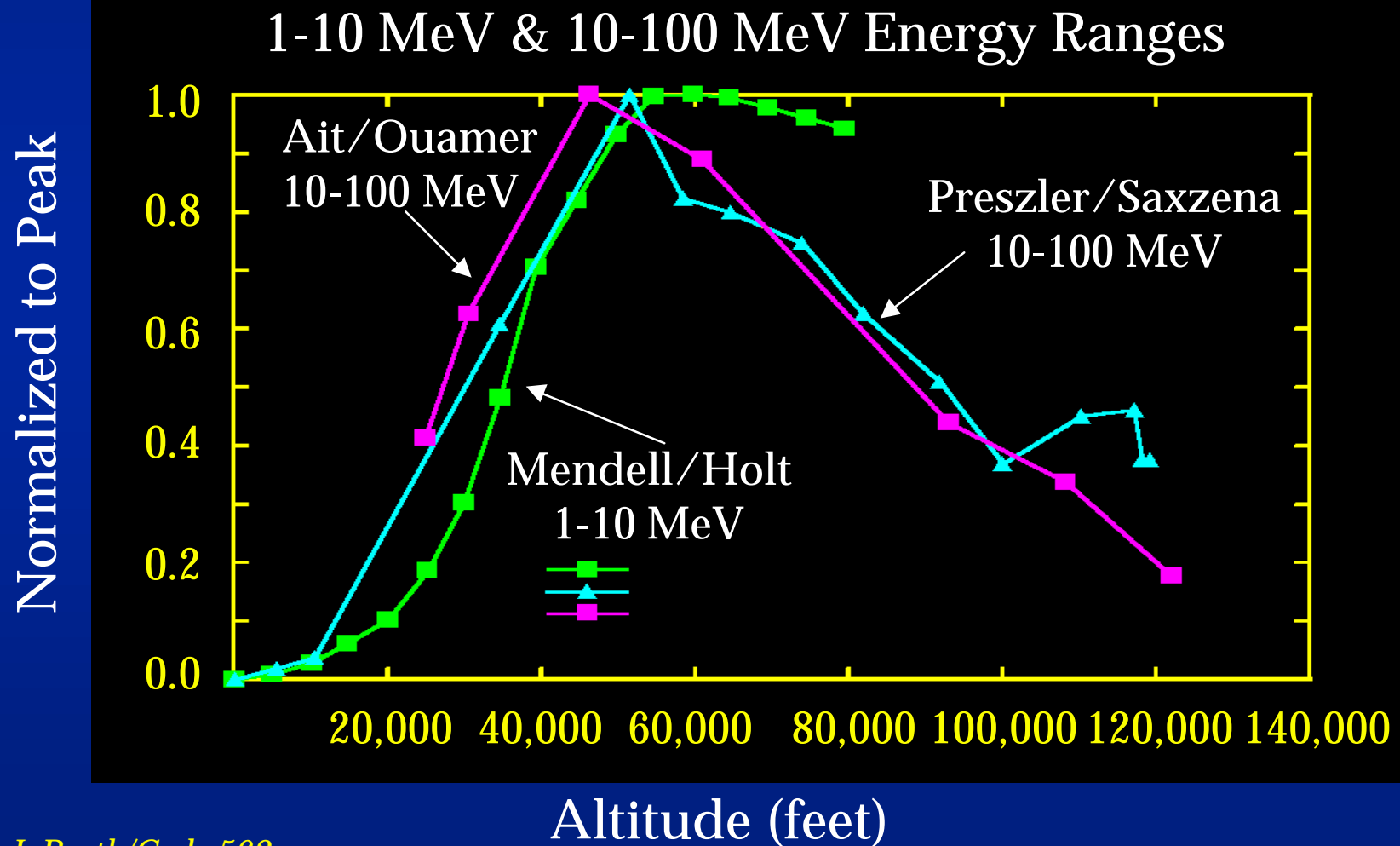
← ~ 35,000 feet
10,000 m

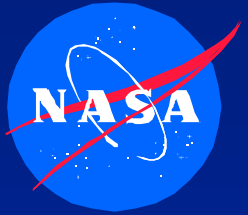
Aircraft

Ground ~ 1/500 of Peak Flux



Neutron Flux Measurements

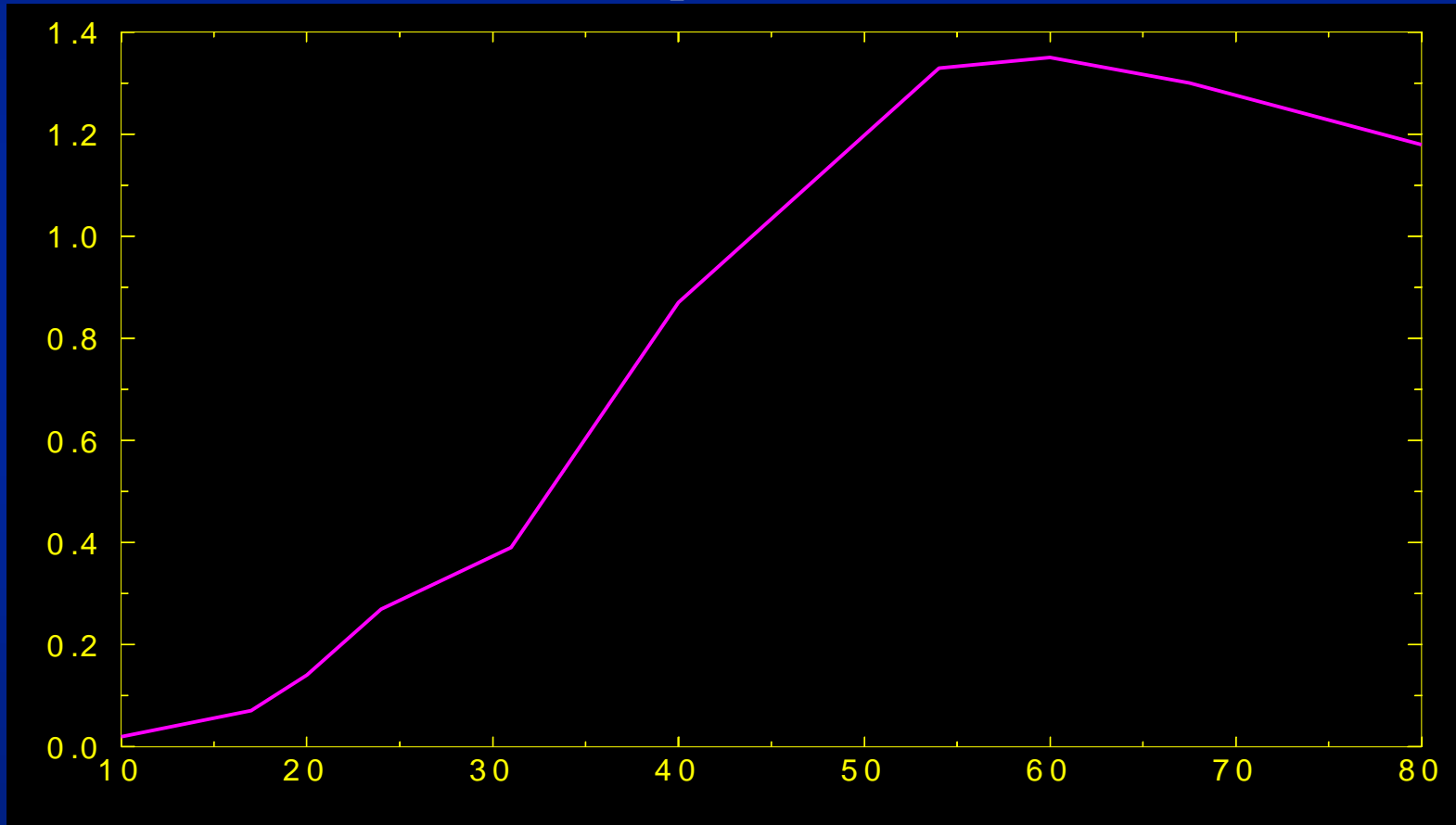




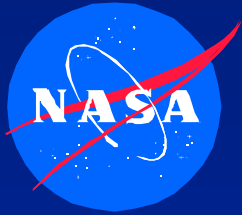
Neutron Models: Flux vs. Altitude

1-10 MeV Atmospheric Neutron Flux

1-10 MeV Neutron Flux ($n/cm^2/s$)



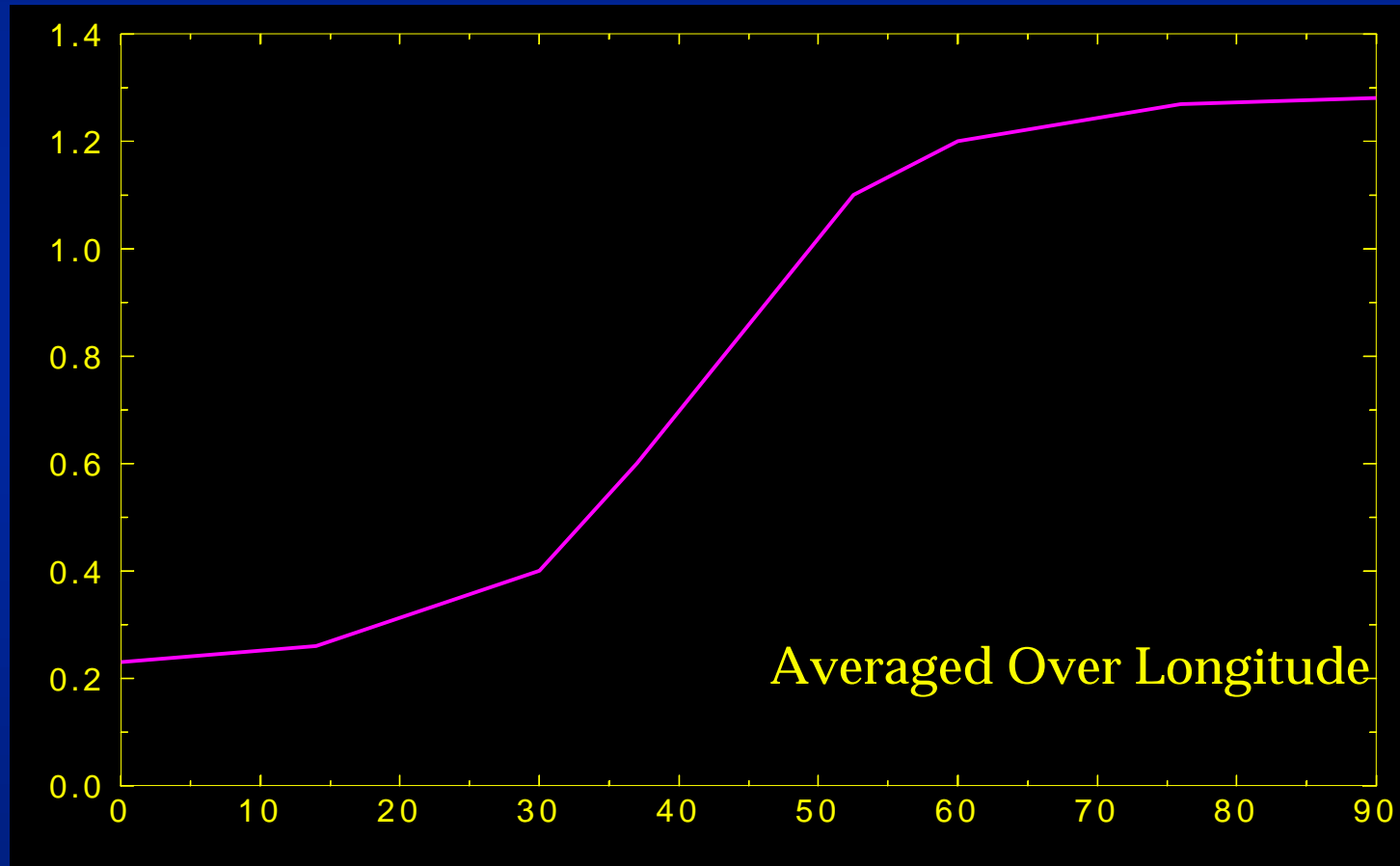
Altitude (Thousands of feet)



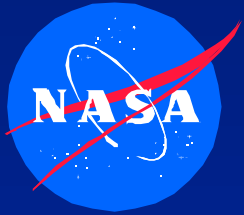
Neutron Model: Flux vs. Latitude

1-10 MeV Atmospheric Neutron Flux

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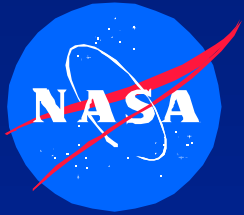


Latitude (deg N)



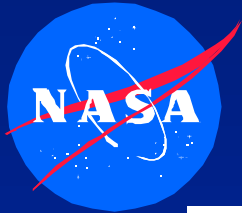
Terrestrial Radiation Sources

- ◆ Man-made
- ◆ Natural Emissions from Earth Materials
 - » Package Contamination
- ◆ Cosmic Rays - Particle Showers

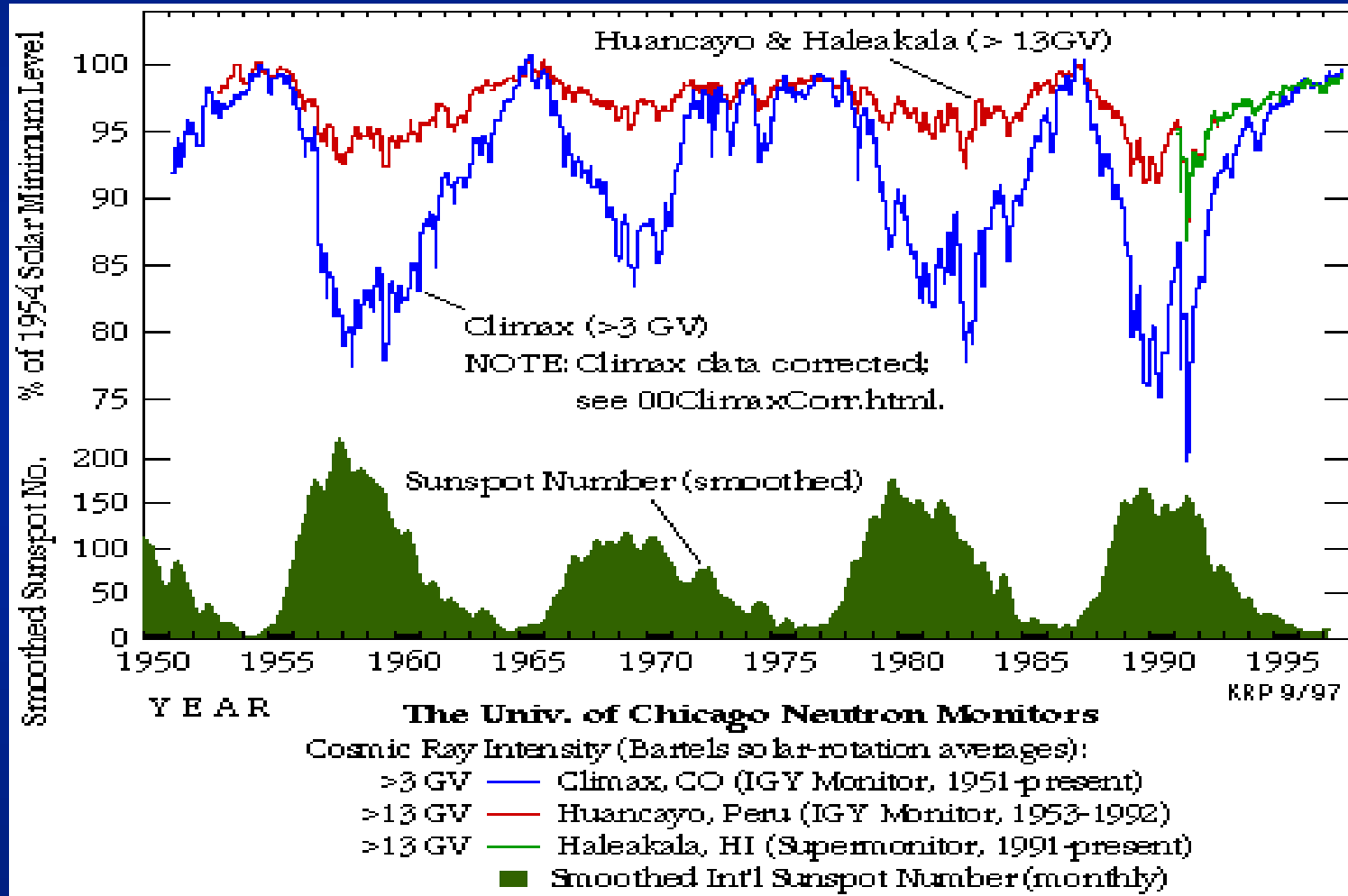


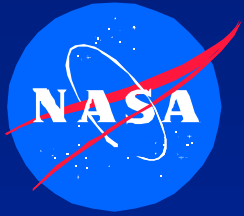
Terrestrial Cosmic Rays

- ◆ Particles That Hit the Earth from Outer Space
- ◆ Need > 1 GeV to Penetrate to Sea Level
- ◆ Fewer Than 1% Are Primary
- ◆ Mostly 3rd to 7th Generation Cascade Particles
- ◆ Search for Cause of Interference on Laboratory Instruments in Early 1900s
 - » Led to Discovery of Cosmic Rays by Hess
- ◆ Induce SEUs: Neutrons + Protons + Pions



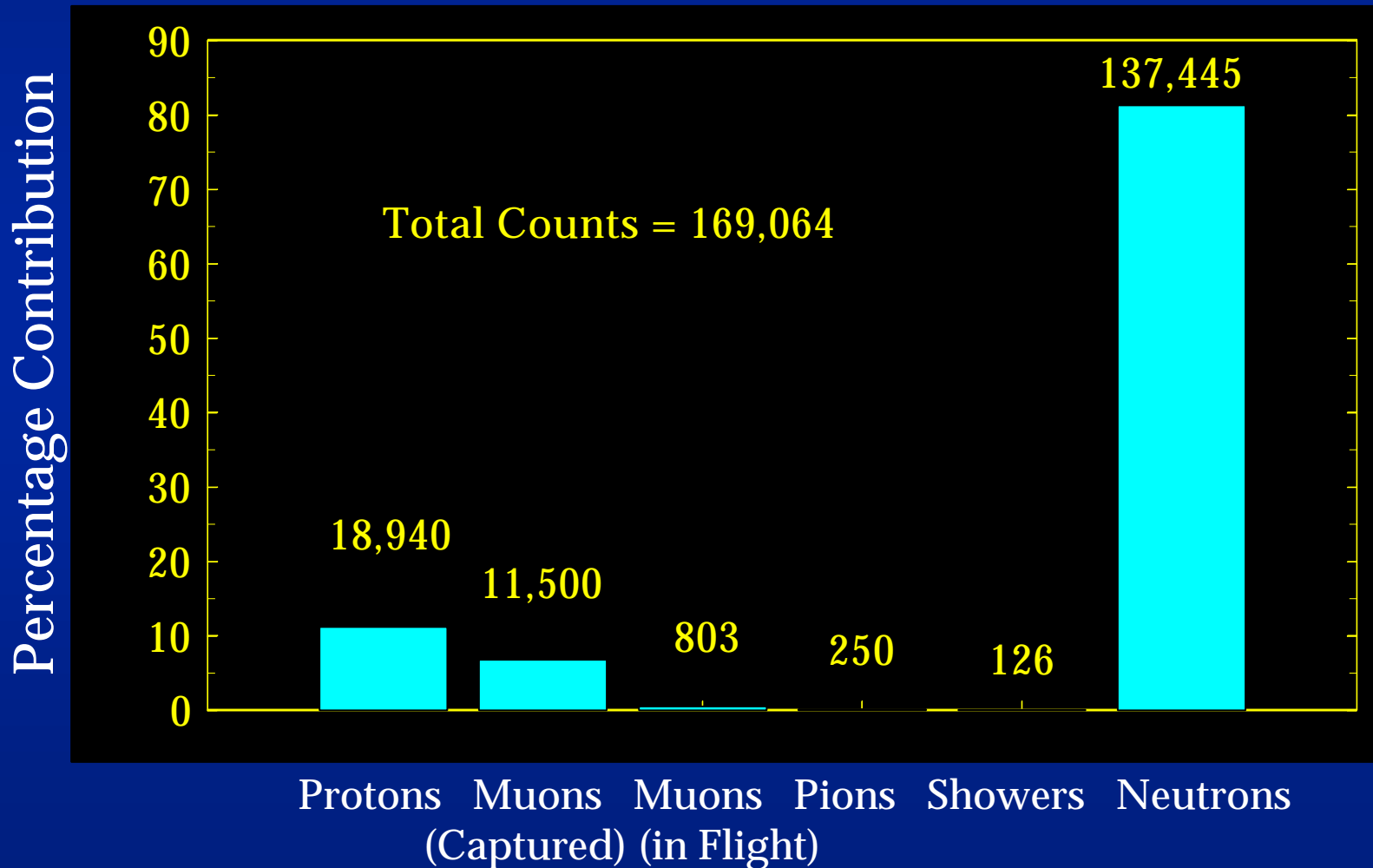
Terrestrial Cosmic Rays

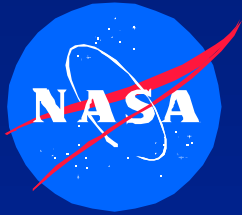




IGY Neutron Monitor Response

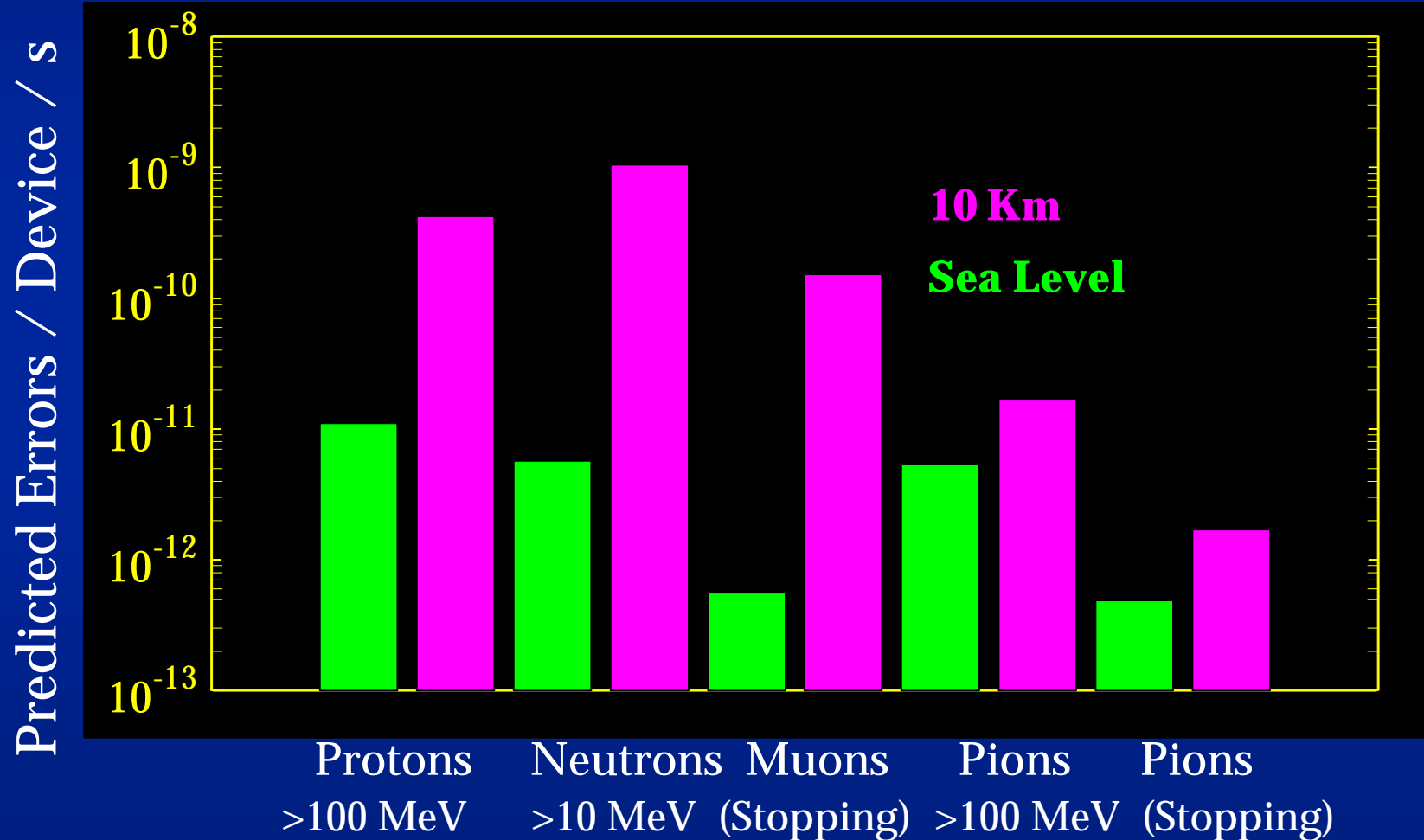
Cosmic Ray Contributions at Sea Level

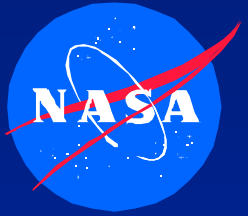




AM9114 4K NMOS Error Rates

Predicted Error Rates at Two Altitudes





Radiation Issues - Three Prime Technical Drivers

- ◆ COTS & Emerging Technologies
 - » More sensitive to radiation
 - » Some devices have new effects
- ◆ COTS - greater uncertainty about radiation hardness
 - » Limited control
 - » Frequent process changes
- ◆ Devices exposed to more radiation on-orbit
 - » Use of composite materials in spacecraft structures
 - » Shrinkage in spacecraft size & weight

Result:

We are using more radiation sensitive components with less protection